

DOES RELATIONALISM ALONE CONTROL GEOMETRODYNAMICS WITH SOURCES?

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Abstract

This paper concerns relational first principles from which the Dirac procedure exhaustively picks out the geometrodynamics corresponding to general relativity as one of a handful of consistent theories. This was accompanied by a number of results and conjectures about matter theories and general features of physics – such as gauge theory, the universal light cone principle of special relativity and the equivalence principle – being likewise picked out. I have previously shown that many of these matter results and conjectures are contingent on further unrelational simplicity assumptions. In this paper, I point out 1) that the exhaustive procedure in these cases with matter fields is slower than it was previously held to be. 2) While the example of equivalence principle violating matter theory that I previously showed how to accommodate on relational premises has a number of pathological features, in this paper I point out that there is another closely related equivalence principle violating theory that also follows from those premises and is less pathological. This example being known as an ‘Einstein–aether theory’, it also serves for 3) illustrating limitations on the conjectured emergence of the universal light cone special relativity principle.

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1 Introduction

1.1 Relationalism

The relational perspective of Barbour [1, 2, 3, 4, 5] implements ideas of Leibniz [6] and Mach [7] (see also [8]) to modern physics. In this approach, one starts with a configuration space \mathcal{Q} of (models of) whole-universe systems. One then adopts two relational postulates.

Configurational relationalism: that certain transformations acting on \mathcal{Q} are physically meaningless. One way [9]¹ of implementing this is to use arbitrary- G -frame-corrected quantities rather than bare \mathcal{Q} configurations, where G is the group of physically meaningless motions. For, despite this augmenting \mathcal{Q} to $\mathcal{Q} \times G$, variation with respect to each adjoined independent auxiliary G -variable produces a constraint which removes one G variable and one redundancy among the \mathcal{Q} variables, so that one ends up on the quotient space \mathcal{Q}/G (the desired reduced configuration space). This is widely a necessity in theoretical physics through working on the various reduced spaces directly often being technically unmanageable, for instance in particle physics theories with its internal gauge group G or in the split spacetime approach to general relativity with its spatial diffeomorphisms.

Temporal relationalism: that there is no meaningful primary notion of time for the universe as a whole. One implementation of temporal relationalism is through using manifestly reparametrization invariant actions that do not rely on any extraneous time-related variables either.

For $\mathcal{Q} = \{n \text{ particle positions}\}$ and G the Euclidean group of translations and rotations, the relational postulates form plausible alternative foundations for a portion of Newtonian mechanics [11, 12, 13] (and admit also a scale-free counterpart for G the Similarity group of translations, rotations and dilatations [10, 14, 12]). The main idea in this paper concerns that (spatially compact without boundary) general relativity can be derived from these postulates in the case in which G is the group of 3-diffeomorphisms. (This derivation also assumes a set of mathematical simplicity postulates and observational checks [4, 15] described in Sec 1.3). This answers a question of Wheeler: “*if one did not know the Einstein–Hamilton–Jacobi equation, how might one hope to derive it straight off from plausible first principles without ever going through the formulation of the Einstein field equations themselves?*” ([16], p 273) (Hojman, Kuchař and Teitelboim [17] had previously provided a distinct answer in which spacetime structure was presupposed; the present answer presupposes less structure than that, being a 3-space rather than split spacetime approach). Finally relational particle models have a number of useful analogue features permitting them to serve as useful [18, 19, 13] toy model analogues for investigations of such as the problem of time in quantum gravity [18, 20].

1.2 General relativity admits a relational formulation

One should first demonstrate that general relativity can indeed be recast as a 3-space approach theory. The Einstein–Hilbert action for the spacetime formulation of general relativity,²

$$I_{\text{GR}}^{\text{EH}}[g_{AB}] = \int d^4x \sqrt{|g|} \mathcal{R} , \quad (1)$$

¹Barbour’s own way of conceptualizing about configurational relationalism (‘best matching’), see e.g. [1, 10] is that, given two configurations, one should be kept fixed and the other should be shuffled around until an identification is found that minimizes its incongruence with the first one. The arbitrary frame method described in the main text here permits the form of the shuffling correction to be derived. Both approaches can be carried out for multiplier or velocity of a cyclic coordinate interpretations of auxiliaries in simple cases (which include all of those covered in this paper).

²Here, g_{AB} is the spacetime metric with determinant g and Ricci scalar \mathcal{R} . h_{ab} is the induced 3-metric on a positive-definite 3-surface Σ (interpreted, for the moment, as a spatial hypersurface within a spacetime), with determinant h , covariant derivative D_a and Ricci scalar R . α is the lapse and β_μ is the shift. $\delta_\beta = \dot{} - \mathcal{L}_\beta$ is the hypersurface derivative, where the dot is $\frac{\partial}{\partial \lambda}$ and \mathcal{L}_β is the Lie derivative with respect to β_a .

when split with respect to a family of spatial hypersurfaces takes the conventional form [21, 22]

$$I_{\text{GR}}^{\text{ADM}}[h_{ab}, \alpha, \beta_a, \dot{h}_{ab}] = \int d\lambda \int d^3x \sqrt{h} \alpha \left\{ \frac{\mathbb{T}_{\text{GR}}^{\text{ADM}}}{4\alpha^2} + R \right\} \quad (2)$$

for

$$\mathbb{T}_{\text{GR}}^{\text{ADM}} = \frac{1}{\sqrt{h}} G^{abcd} \{ \delta_\beta h_{ab} \} \delta_\beta h_{cd} \quad (3)$$

and

$$G^{abcd} = \sqrt{h} \{ h^{ac} h^{bd} - h^{ab} h^{cd} \} \quad (4)$$

the inverse of the DeWitt supermetric [22].

A more useful prototype 3-space approach action [1] can be formed by Baierlein, Sharp and Wheeler's [23] procedure: solve the α -multiplier equation $R - \mathbb{T}_{\text{GR}}^{\text{ADM}}/4\alpha^2 = 0$ for α itself, $\alpha = \frac{1}{2} \sqrt{\mathbb{T}_{\text{GR}}^{\text{ADM}}/R}$, and then use this to algebraically eliminate α from the Arnowitt–Deser–Misner Lagrangian. Thus one obtains

$$I_{\text{GR}}^{\text{BSW}}[h_{ab}, \beta_a, \dot{h}_{ab}] = \int d\lambda \int d^3x \sqrt{h} \sqrt{R \mathbb{T}_{\text{GR}}^{\text{ADM}}} . \quad (5)$$

This is not quite reparametrization invariant because the shift is considered to be a coordinate for the purposes of variation. However, the Arnowitt–Deser–Misner split can be replaced by a split in terms of an instant variable \mathbb{I} (such that $\alpha = \mathbb{I}$) and a grid variable F_a (such that $\beta_a = \dot{F}_a$, which is an example of a frame variable) at the pre-variational level if one takes into careful account that the auxiliary variables should be varied with free end spatial hypersurfaces [24].³ Then one has an action

$$I_{\text{GR}}^{\mathbb{A}}[h_{ab}, \dot{h}_{ab}, \dot{F}_a, \mathbb{I}] = \int d\lambda \int d^3x \sqrt{h} \mathbb{I} \left\{ \frac{\mathbb{T}_{\text{GR}}^{\mathbb{A}}}{4\mathbb{I}^2} + R \right\} . \quad (6)$$

for

$$\mathbb{T}_{\text{GR}}^{\mathbb{A}} = \frac{1}{\sqrt{h}} G^{abcd} \{ \delta_{\dot{F}} \} h_{ab} \delta_{\dot{F}} h_{cd} . \quad (7)$$

Then performing Routhian reduction on this to eliminate \mathbb{I} works out exactly the same as Baierlein–Sharp–Wheeler multiplier elimination, giving

$$I_{\text{GR}}^{\mathbb{A}'}[h_{ab}, \dot{h}_{ab}, \dot{F}_a] = \int d\lambda \int d^3x \sqrt{h} \sqrt{R \mathbb{T}_{\text{GR}}^{\mathbb{A}}} . \quad (8)$$

This may now be taken as a starting point as done in [4, 25, 9] that implements the relational principles, in which case I use the notation $\&_{\dot{F}}$ for arbitrary G frame corrected derivative, here for G the 3-diffeomorphisms on Σ :

$$I_{\text{GR}}^{\text{TSA}}[h_{ab}, \dot{h}_{ab}, \dot{F}_a] = \int d\lambda \int d^3x \sqrt{h} \sqrt{R \mathbb{T}_{\text{GR}}^{\text{TSA}}[h_{ab}, \dot{h}_{ab}, \dot{F}_a]} , \quad \mathbb{T}_{\text{GR}}^{\text{TSA}} = \{ h^{ac} h^{bd} - h^{ab} h^{cd} \} \{ \&_{\dot{F}} h_{ab} \} \&_{\dot{F}} h_{cd} , \quad (9)$$

rather than the hypersurface derivative notation $\delta_{\dot{F}}$ that presupposes spacetime. Of course, in the present case, spacetime is nevertheless recovered.

³See [10, 25, 26, 27] for earlier and further discussion of these variational methods.

1.3 The ‘relativity without relativity’ result

Suppose next that one presupposes less structure: just 3-space notions rather than ‘3-space within spacetime’ notions. Does general relativity then emerge? Does it emerge alone? One goes about investigating these questions using the *Dirac procedure* [28]. This involves taking the constraints that arise purely from the form of the Lagrangian without any variation (primary constraints) and those that have arisen so far by the variational process (secondary constraints), and demanding that these be propagated by the theory’s evolution equations. This can lead to new constraints, in which case the Dirac procedure is applied again to these. Now, as each new constraint uses up some degrees of freedom (usually per space point in the present field-theoretic context) and the trial system has a finite amount of these, if the Dirac procedure runs through enough iterations, it uses up at least as many degrees of freedom as the trial theory had to start off with (see e.g. [29, 30]. In this case, the trial theory has been demonstrated to be undesirable in being inconsistent (less than no degrees of freedom left), trivial (no degrees of freedom left) or undersized (e.g. a few global degrees of freedom alone could survive due to the shapes of the restrictions caused by the constraints, see e.g. [9]). Then the only remaining way out is to restrict the trial theory by allowing some of the constraints to dictate how some of its hitherto free non-variational parameters should be fixed, and so one is exhaustively removing a number of the trial options. Thus the Dirac procedure lends itself to proofs by exhaustion.

By this method the ‘relativity without relativity’ result arises: if one does not presuppose general relativity but rather to start with a wide class of reparametrization-invariant actions built out of good 3-d space objects in accord with the relational principles [4, 31, 32, 9, 15], general relativity emerges. More concretely the input trial ansätze are

$$\mathbb{T}_{\text{grav(trial)}} = \frac{1}{\sqrt{\hbar}Y} G^{abcd}(W) \{ \&_{\mathbb{F}} h_{ab} \} \&_{\mathbb{F}} h_{cd} , \quad (10)$$

for the gravitational kinetic term that is **homogeneous quadratic in the velocities**, where

$$G^{ijkl}(W) \equiv \sqrt{\hbar} \{ h^{ik} h^{jl} - W h^{ij} h^{kl} \} , \quad W \neq \frac{1}{3} , \quad (11)$$

is the inverse of the most general (**invertible, ultralocal**) supermetric

$$G_{abcd}(X) = \frac{1}{\sqrt{\hbar}} \left\{ h_{ac} h_{bd} - \frac{X}{2} h_{ab} h_{cd} \right\} , \quad X = \frac{2W}{3W - 1} , \quad (12)$$

and

$$\mathbb{V}_{\text{grav(trial)}} = A + BR \quad (13)$$

for the gravitational potential term. This is **second-order in spatial derivatives**. The **local square root** action is then

$$\mathbb{I}_{\text{grav(trial)}}[h_{ab}, \dot{h}_{ab}, \dot{\mathbb{F}}_i] = \int d\lambda \int d^3x \sqrt{\hbar} \sqrt{\mathbb{V}_{\text{grav(trial)}} \mathbb{T}_{\text{grav(trial)}}} . \quad (14)$$

[I use bold font to denote what assumptions are being made; all of the assumptions in this Subsection are *mathematical simplicity postulates* rather than deep physical principles.]

Then, setting $\dot{\mathbb{M}}$ to be the emergent quantity $\frac{1}{2} \sqrt{\mathbb{T}_{\text{grav(trial)}}^{\text{TSA}} / \mathbb{V}_{\text{grav(trial)}}}$, the gravitational momenta are

$$\pi^{ab} \equiv \frac{\partial \mathbb{L}}{\partial \dot{h}_{ab}} = \frac{\sqrt{\hbar}Y}{2\dot{\mathbb{M}}} G^{abcd}(W) \&_{\mathbb{F}} h_{cd} , \quad (15)$$

which are related by a primary constraint

$$\mathcal{H}_{\text{grav(trial)}} \equiv Y G_{abcd}(X) \pi^{ab} \pi^{cd} - \sqrt{\hbar} \{ A + BR \} = 0 . \quad (16)$$

Additionally, variation with respect to F_a leads to a secondary constraint that is the usual momentum constraint

$$\mathcal{H}_a = D_b \pi_a^b = 0 \quad (17)$$

thereby ensuring that the physical content of the theory is in the shape of the 3-geometry and not in the coordinate grid painted on it. The propagation of $\mathcal{H}_{\text{grav(trial)}}$ then gives [31, 9]

$$\dot{\mathcal{H}}_{\text{grav(trial)}} \approx \frac{2}{M} \{X - 1\} B Y D_i \{M^2 D^i \pi\}, \quad (18)$$

[where \approx is Dirac's notion of weak equality, i.e. equality up to (already-known) constraints].

From this, the main output is the 'relativity without relativity' result that the Hamiltonian constraint propagates if the coefficient in the supermetric takes the DeWitt value $X = 1 = W$. In this case, embeddability of the 3-space into spacetime is recovered. This in no way determines whether the emergent spacetime's signature is Lorentzian ($B = -1$) or Euclidean ($B = 1$): that is to be put in by hand.

1.4 General relativity as geometrodynamics does not arise alone in the 3-space approach

As it has 3 further factors [4, 31, 33, 15], (18) can vanish in 3 other ways.

1) $B = 0$ gives strong or 'Carrollian' gravity options regardless of whether $W = 1$ or not. The $W = 1$ case is the strong-coupled limit of general relativity [34], which is a regime in which distinct points are causally disconnected by their null cones being squeezed into lines. This is relevant as an approximation to general relativity near singularities. While, for $W \neq 1$, it is a similar limit of scalar-tensor theories [31]. In fact, all of these $B = 0$ options exist in two different forms: one without a momentum constraint which thus are temporally but not spatially relational *metrodynamics* and one with a momentum constraint which are other consistent theories of geometrodynamics different to that obtained from decomposing the spacetime formulation of general relativity.

2) $Y = 0$, gives 'Galilean' theories. Here, the null cones are squashed into planes and there is no gravitational kinetic term. Strictly speaking, for this option to arise, one should start with the Hamiltonian version of the 'Galilean' theory (as its degeneracy leads to there being no corresponding Lagrangian).

3) $\pi = 0$ or $\pi/\sqrt{\hbar} = \text{constant}$ preferred slicing conditions make the fourth factor vanish. This gives rise to alternative theories of conformal gravity [25] and to a *derivation of general relativity, alongside the conformal method of treating its initial-value problem* [35], *from a relational perspective* [25, 26, 27]. These theories can be recast by restarting with an enlarged irrelevant group G that consists of the 3-diffeomorphisms together with some group of conformal transformations.

1.5 Inclusion of matter in the 3-space approach

The second theme of the 3-space approach papers concerns the inclusion of fundamental matter. This is important for the 3-space approach, both as a robustness test for the axiomatization and to establish whether special relativity and the equivalence principle are emergent or require presupposition in this approach.

The robustness test is passed: using first constructive techniques [4, 29] and then Kuchař's [36] split spacetime framework techniques ([32], see also Sec 3), all of minimally-coupled scalars, electromagnetism, Yang-Mills theory, Dirac theory, and all the associated gauge theories were found to be admitted by the 3-space approach. There were some claims as regards well-known matter field types and physical principles being picked out. For example, it was claimed that

- 1) That electromagnetism and Yang-Mills theory are uniquely picked out [4, 29].
- 2) That minimally coupled scalars and 1-forms share null cones among themselves (which is evidence toward the emergence of the special relativity principle). This is through each being forced to share the gravitational null cone.
- 3) That the equivalence principle is emergent rather than assumed.

These were based on the simplicity assumptions of **matter kinetic terms homogeneous quadratic in their velocities, no metric-matter kinetic cross terms, no matter dependence in the kinetic metric.**

However, the split spacetime framework and allied techniques proved powerful enough to include massive (and other) vector fields [9, 33, 15] if these simplicity postulates are weakened in various ways, showing that the latter claims are partly based on mere simplicities that have nothing to do with relationalism. Hence 1) is false. Furthermore, this paper demonstrates that 2) and 3) are false. Essentially the split spacetime framework suggests further terms for the kinetic and potential ansatz with the inclusion of which further consistent theories can be cast in 3-space approach form. Overall, the relational postulates do not pick out the fields of nature, they include a wider range of fields.

A further new point I make in the present paper is that even within the simplicity postulates assumed, the claims were based on calculations that have two further tacit simplicity assumptions, without which the exhaustion rate would be slower than it was held to be.

1) **Linear combination constraint preclusion.** In the original calculations, constraints arising as linear combinations of terms with different a priori free parameters were not considered to be a possibility. However, there is no good theoretical reason to preclude such constraints from arising.

2) **Second class constraint preclusion.** The original calculations' counting implicitly assumed that all constraints arising were first-class as regards how many degrees of freedom they used up.⁴

While preclusion 2) is a brief and mathematically well defined simplicity postulate, it is highly restrictive, e.g. it does not cover the usual presentation of the phenomenologically useful massive vector field. This sort of restriction makes it very desirable from a theoretical perspective to uplift this simplicity. One way to proceed as regards 2) (which is simple and rigorous, although it is clearly not the most efficient) is to only assume that each constraint uses up at least one degree of freedom.

1) and 2) are clearly then capable of increasing the number of iterations required before a theory is shown to be inconsistent. In particular, for a set of interacting vector fields, weakening 1) costs one the capacity to produce internal index valued constraints at each step, meaning that one can no longer work for 'arbitrary' gauge group.⁵ All that is known now then is that for *fairly small* gauge groups the calculation excludes alternatives, the calculation remaining unfinished for larger gauge groups. Thankfully, the gauge groups that have been found to explain experimental particle physics are not too large... (One can thus work furthermore case-by-case for larger groups required for more speculative theories of particle physics such as grand unified theories, while one should also not rule out that some new efficiency trick could be found so as to recover the result for an 'arbitrary group'). One could likewise work case-by-case so as to safeguard other previous claims such as those about higher potential derivatives in vacuo in [4].

1.6 Outline of the rest of this paper

The constructive workings of [4, 29], all of which assume homogeneous quadratic kinetic terms with no metric–matter cross terms or matter field dependence in the kinetic metric, suffice as an arena in which to investigate the local emergence of special relativity (Sec 2), at least for simple 3-space approach theories. In Sec 3, I recollect (and add to) arguments against the assertion (p 3217 of [4]) that in the 3-space approach “*self-consistency requires that any 3-vector field must satisfy ... the equivalence principle*”. These arguments involve casting scalar(–vector)–tensor theories into 3-space approach form to act as counterexamples. I add further to these arguments in Sec 4 by constructing a unit vector tensor theory in 3-space approach form that is free of some pathologies common to vector–tensor theories and is both equivalence principle violating and special relativity violating in the sense that it has more than one distinct finite fundamental propagation speed. Hence relationalism alone does not locally imply the special relativity principle.

⁴A constraint is *first-class* if its Poisson brackets with all the other constraints close, and *second-class* otherwise. First-class constraints use up two degrees of freedom each while second-class ones use up just one. However one does not know before the Dirac process terminates whether a constraint is first or second class – do its Poisson brackets with as yet unfound constraints from further along the Dirac progress close? Thus one cannot argue for emergent constraints to use up two degrees of freedom each (at least until the Dirac process has terminated and one has evaluated all those Poisson brackets).

⁵'Arbitrary' here is subject to the (usual) requirement of being a direct sum of compact simple and U(1) Lie subalgebras so that the Gell-Mann–Glashow theorem applies [37].

2 The position hitherto about the emergence of special relativity

On p4 of [28], Dirac explains that he uses actions so that relativity and gauge symmetry can be straightforwardly incorporated from the start. This is done by constructing one's action out of quantities that are Poincaré invariant for special relativity, diffeomorphism-invariant for general relativity, U(1) gauge invariant for electromagnetic theory, and so on. The 3-space approach is in a sense a reverse of this: neither spacetime structure nor its locally special relativistic element are presupposed and it is shown that most alternatives to this are inconsistent. The way in which the early 3-space approach papers [4, 29] include a range of standard bosonic matter fields minimally coupled to general relativity is a sufficient arena to investigate whether and how special relativity locally emerges in the 3-space approach. These papers make the homogeneously-quadratic kinetic ansatz $\mathbb{T} = \mathbb{T}_\Psi + \mathbb{T}_{\text{grav}(\text{trial})}$, where the matter fields Ψ_Δ have kinetic term

$$\mathbb{T}_\Psi = G^{\Gamma\Delta}(h_{ij})\{\&_{\hat{F}}\Psi_\Gamma\}\&_{\hat{F}}\Psi_\Delta, \quad (19)$$

potential term denoted by \mathbb{U}_Ψ and momenta denoted by Π^Δ .

Then the implementation of temporal relationalism by reparametrization invariance leads to a Hamiltonian-type constraint

$$\mathcal{H}_{\text{grav-}\Psi(\text{trial})} \equiv \sqrt{\hbar}\{A + BR + \mathbb{U}_\Psi\} - YG_{abcd}(X)\pi^{ab}\pi^{cd} + \frac{G_{\Gamma\Delta}\Pi^\Gamma\Pi^\Delta}{\sqrt{\hbar}} = 0. \quad (20)$$

Applying Dirac's procedure and assuming that \mathbb{U}_Ψ at worst depends on connections (rather than their derivatives, which is true for the range of fields in question), the propagation of $\mathcal{H}_{\text{grav},\Psi(\text{trial})}$ gives

$$\begin{aligned} \dot{\mathcal{H}}_{\text{grav-}\Psi(\text{trial})} \approx & \frac{2}{\mathbb{M}}D^a \left\{ \dot{\mathbb{M}}^2 \left\{ Y \left\{ B \left\{ D^b\pi_{ab} + \{X-1\}D_a\pi \right\} + \right. \right. \\ & \left. \left. \left\{ \pi_{ij} - \frac{X}{2}\pi h_{ij} \right\} \left\{ \frac{\partial\mathbb{U}_\Psi}{\partial\Gamma^c_{ia}}h^{cj} - \frac{1}{2}\frac{\partial\mathbb{U}_\Psi}{\partial\Gamma^c_{ij}}h^{ac} \right\} \right\} + G_{\Gamma\Delta}\Pi^\Gamma \frac{\partial\mathbb{U}_\Psi}{\partial(\partial_a\Psi_\Delta)} \right\}, \end{aligned} \quad (21)$$

which is just an extension of (18) to include some matter fields. The terms in (21) are then required to vanish for consistency. This can occur according to various options, each of which imposes restrictions on $\mathcal{H}_{\text{grav-}\Psi(\text{trial})}$. Furthermore, these options turn out to be very much connected to those encountered in the usual development of special relativity.

There is now a three-pronged fork in the choice of a universal transformation law in setting up special relativity. Two prongs are the Galilean or Lorentzian fork that Einstein faced (infinite or finite universal maximum propagation speed c). The third prong is the Carrollian option $c = 0$. This last option occurs above through setting $B = 0$. The vanishing of the other factors is attained by 1) declaring that \mathbb{U}_Ψ cannot contain connections. 2) It is then 'natural' for \mathbb{U}_Ψ not to depend on $\partial_a\Psi_\Delta$ either (ultralocality in Ψ_Δ), whereby the last term is removed. Of course, we have good reasons to believe nature does not have $c = 0$, but what this option does lead to is alternative dynamical theories of geometry to the usual general relativistic geometrodynamics. Some are spatially relational and some are not. This is an interesting fact from a broader perspective: it issues a challenge to why the 3-space approach insists on geometrodynamical theories since metrodynamical theories are also possible. But what happens in the general relativity option is that the momentum constraint is an integrability condition [38, 31] so one is stuck with geometrodynamics whether one likes it or not.

One could also enforce consistency above by the 'Galilean' strategy of choosing $Y = 0$. This removes all but the last term. It would seem natural to take this in combination with $\Pi^\Delta = 0$, whereupon the fields are not dynamical. Moreover this does not completely trivialize the matter fields since they would then obey analogues of Poisson's law, or Ampère's, and these are capable of governing a wide variety of complicated patterns. Thus one arrives at an entirely nondynamical 'Galilean' world. In vacuo, this possibility cannot be obtained from a Baierlein-Sharp-Wheeler-type Lagrangian (the kinetic factor is badly behaved) but the Hamiltonian description of the theory is unproblematic. Of course, the Hamiltonian constraint is now no longer quadratic in the momenta:

$$\mathcal{H}_{\text{grav-}\Psi(\text{trial})}(Y = 0) = A + BR + \mathbb{U}_\Psi = 0. \quad (22)$$

This option is not of interest if the objective is to find *dynamical* theories. Nevertheless, this option is a logical possibility, and serves to highlight how close parallels to the options encountered in the development of special relativity arise within the 3-space approach.

There is also a combined locally Lorentzian physics and spacetime structure strategy as follows. The signature is to be set by hand (one could just as well have any other nondegenerate signature for the argument below). Take (21) and introduce the concept of a gravity–matter momentum constraint $\mathcal{H}_{\text{grav-}\Psi(\text{trial})}^a$ by using $0 = -\frac{1}{2}\mathcal{H}_{\Psi}^a + \frac{1}{2}\mathcal{H}_{\Psi}^a$ and refactoring:

$$\begin{aligned} \dot{\mathcal{H}}_{\text{grav-}\Psi(\text{trial})} \approx & \frac{2D^a}{\dot{M}} \left\{ \dot{M}^2 \left\{ Y \left\{ B \left\{ \left\{ \underline{\underline{D^b \pi_{ab} - \frac{1}{2} \left[\Pi^\Delta \frac{\delta \mathcal{L}_{\dot{F}} \Psi_\Delta}{\delta \dot{F}^a}} \right]}} \right\} + \frac{1}{2} \left[\Pi^\Delta \frac{\delta \mathcal{L}_{\dot{F}} \Psi_\Delta}{\delta \dot{F}^a} \right] \right\} \right\} + \underline{\underline{G_{\Gamma\Delta} \Pi^\Delta \frac{\partial U_\Psi}{\partial (\partial_a \Psi_\Delta)}}}} \right. \\ & \left. + \underline{\underline{Y B \{ X - 1 \} D_a \pi}} + Y \left\{ \underline{\underline{\pi_{ij} - \frac{X}{2} \pi h_{ij}}} \right\} \left\{ \underline{\underline{\frac{\partial U_\Psi}{\partial \Gamma_{ia}^c} h^{cj} - \frac{1}{2} \frac{\partial U_\Psi}{\partial \Gamma_{ij}^c} h^{ac}}} \right\} \right\} , \end{aligned} \quad (23)$$

so that the first two underlined terms are then proportional to $\mathcal{H}_{\text{grav-}\Psi(\text{trial})}^a$.⁶ In the ‘orthodox general covariance option’, the third and fourth underlined terms cancel, amounting to the enforcement of a universal null cone. This requires supplementing by some means of discarding the fifth underlined term. Here, one can furthermore *choose* the orthodox option $X = 1$: the recovery of embeddability into spacetime corresponding to general relativity (the ‘relativity without relativity’ result), or, *choose* the alternative preferred-slicing worlds of $D_a \pi = 0$ which are governed by conformal mathematics. As both of these options are valid, the recovery of locally-Lorentzian physics does not occur solely in generally-covariant theories. The connection terms (sixth underlined term) must also be discarded, but the Dirac procedure does this automatically for the given ansätze.

Thus in the 3-space approach, locally-Lorentzian general relativistic spacetime arises as one option; other permitted options include Carrollian worlds, Galilean worlds and locally-Lorentzian preferred slicing worlds. These alternatives all lack some of the features of generally relativistic spacetime. [15] went on to talk about hybrids of the above. The ultralocal and nondynamical strategies for dealing with the last term in (21) are available in *all* the above options. So as things stand, one does derive that gravitation enforces a *unique finite* propagation speed, but the possibility of coexisting with fields with infinite and zero propagation speeds is not precluded by consistency, although it does read to undersized solution spaces.⁷ And of making the fourth underlined term vanish algebraically along the lines of parallel E and B in Poynting vector. But none of these situations ruin the emergence of the special relativity principle in the sense that: any adjoined zero-momentum Galilean fields cannot propagate so that it does not matter that their propagation speed is in principle infinite, while adjoined Carrollian fields are precluded from propagating information away from any point by their ultralocal nature, and the parallel E and B field situation is also well-known to preclude the associated propagation (mutual orthogonality in the E and B fields causing each other to continue to oscillate).

However, we shall see in Sec 4 that the above fork breaks down for more complicated matter.

⁶ $\frac{\delta A}{\delta B}$ denotes the functional derivative, and the special brackets [] delineate the factors over which the implied integration by parts is applicable.

⁷ That such a dilemma exists was simply overlooked in [4, 29] papers since it was claimed that these ultralocal and nondynamical strategies only lead to trivial theories, in the latter case by counting arguments. Unfortunately, inspection of this triviality reveals it to mean ‘less complicated than in conventional Lorentzian theories’ rather than ‘devoid of mathematical solutions’. In particular, the counting argument is insufficient in not taking into account the geometry of the restrictions on the solution space.

3 The Position hitherto on the equivalence principle in the 3-space approach

3.1 Equivalence principle violations at the level of the action

While this study started with partial evidence for the equivalence principle being emergent in the 3-space approach [4, 9], it then suffered the setback of counterexamples as more complete potential ansätze were devised. The counterexamples to date have, however, suffered from certain limitations. In this paper I extend the counterexamples to cases for which these limitations do not occur. I should first describe some symptoms at the level of action principles of whether a theory obeys or violates the equivalence principle. Coordinates can be provided at each particular point p such that the metric connection vanishes at p , so there is no obstruction in passage to the local special relativity form for curved spacetime matter field equations⁸ that contain no worse than metric connection. However, the curvature tensor is an obstruction to such a passage if the field equations contain derivatives of the metric connection. Thus theories in which the matter terms contribute additional such terms are equivalence principle violating. One way in which derivatives of the metric connection in the field equations can arise from actions is through there already being such derivatives in the action e.g. in curvature–matter coupling terms. A second way is from integration by parts during the variational working causing mere metric connections in the action to end up as derivatives of metric connections in the field equations.

3.2 The split spacetime framework

Rather than the previous sections’ exhaustive Dirac procedure, this section requires the split spacetime framework, which does presuppose the general relativistic notion of spacetime. The point of this is that there is then a systematic treatment of Kuchař [36] by which the spacetime formulation of specific consistent matter theories can be recast in split spacetime framework form. It was using the split spacetime framework [32, 9] that many matter theories were found to admit formulations that conform to the 3-space approach’s relational principles (Sec 1.5). I next provide (as a new result) the variant of the split spacetime framework that is in terms of instant-grid variables for the case relevant here: a 1-form matter field.

One is presupposing that one has a hypersurface Σ within a spacetime M . n_A is the normal to Σ and e_A^a the projector onto this hypersurface. Then it is meaningful to decompose each matter field into perpendicular and tangential parts with respect to Σ . In the case of the 1-form,

$$A_A = n_A A_\perp + e_A^a A_a . \quad (24)$$

Hypersurfaces can be re-gridded and deformed. Re-gridding kinematics involves Lie derivatives with respect to \dot{F}_a ; these appear as corrections to the velocities so that these feature in the action as ‘hypersurface derivatives’ rather than as ‘bare velocities’. As regards deformations, the arbitrary deformation of a hypersurface near a point p splits into a *translation part* such that

$$\dot{I}(p) \neq 0 , \quad \{\dot{I}, a\}(p) = 0 \quad (25)$$

and a *tilt part* such that

$$\dot{I}(p) = 0 , \quad \{\dot{I}, a\}(p) \neq 0 . \quad (26)$$

The translation piece further splits into a translation on a background spacetime piece and a *derivative coupling* piece which alters the nature of the background spacetime. Furthermore, the re-gridding, tilt and derivative-coupling kinematics pieces that arise within this spacetime-presupposing framework are *universal*: they depend solely on the rank of the tensor matter field rather than on any details of that particular field. It then so happens that two of these bear a tight relationship with what is needed to implement the configurational relationalism and temporal relationalism postulates.

⁸N.B. that the gravitational field equations are given a special separate status in the equivalence principle (‘all the laws of physics bar gravity’) and thus do not interfere with the logic of this.

- 1) Use of the arbitrary 3-diffeomorphism frame is none other than re-gridding.
- 2) The absence of tilt terms is a guarantee of an algebraic Routhian reduction procedure. Thereby an extraneous time variable free action can be obtained, at least in principle.⁹ Tilts can however be removed from at least some actions by such as integration by parts or redefining field variables.

On the other hand, the derivative coupling universal feature is related to the equivalence principle, in that theories in which derivative coupling features in the split action are equivalence principle violating. Absence of derivative coupling is termed the **geometrodynamical equivalence principle** in [17]. It corresponds to the no metric–matter cross–term and no matter field dependence in the kinetic metric. Thus these particular mathematical simplicity postulates additionally have physical significance. [But demanding that these hold amounts to imposing aspects of the equivalence principle by hand, so that one can no longer claim that the equivalence principle is emergent in the 3-space approach.]

This paper requires the split spacetime form of the derivatives of a 1-form, which is also a good illustration of re-gridding, tilt and derivative coupling terms (the last of these are those terms that involve the extrinsic curvature

$$K_{ab} = -\frac{1}{2\dot{\mathbf{I}}}\delta_{\dot{\mathbf{F}}}h_{ab} \quad (27)$$

of the hypersurface).

$$\nabla_b A_\perp = D_b A_\perp - K_{bc} A^c, \quad (28)$$

$$\dot{\mathbf{I}}\nabla_\perp A_a = -\delta_{\dot{\mathbf{F}}}A_a - \dot{\mathbf{I}}K_{ab}A^b - A_\perp\partial_a\dot{\mathbf{I}}, \quad (29)$$

$$\nabla_b A_a = D_b A_a - A_\perp K_{ab}, \quad (30)$$

$$\dot{\mathbf{I}}\nabla_\perp A_\perp = -\delta_{\dot{\mathbf{F}}}A_\perp - A^a\partial_a\dot{\mathbf{I}}, \quad (31)$$

Intuitively, these relations come about because spacetime derivatives are not equal to spatial derivatives as the former have extra connection components, which this scheme interprets geometrically from the perspective of the hypersurface.

3.3 Scalar–tensor theories

The 3-space approach counterexample to date of the first type is Brans–Dicke theory [39] (see also [40] for a canonical treatment). While this was included in [4] by casting the theory in the Einstein frame¹⁰ However, this transformation does away with the equivalence principle violation, so it is more instructive for the present context to work in Brans–Dicke theory’s usual Jordan frame. For this, the spacetime action is

$$\mathfrak{I}_{\text{BD}}[g_{AB}, \chi] = \int d^4x \sqrt{|g|} e^{-\chi/2} \{ \mathcal{R} - \omega \partial_A \chi \partial^A \chi \}. \quad (32)$$

The subsequent split spacetime action has a kinetic term proportional to

$$\left\{ h^{ac}h^{bd} - \frac{X-2}{3X-4}h^{ab}h^{cd} \right\} \delta_{\dot{\mathbf{F}}}h_{ab}\delta_{\dot{\mathbf{F}}}h_{cd} + \frac{4}{3X-4}h^{ab}\delta_{\dot{\mathbf{F}}}h_{ab}\delta_{\dot{\mathbf{F}}}\chi + \frac{3X-2}{(3X-4)(X-1)}\delta_{\dot{\mathbf{F}}}\chi\delta_{\dot{\mathbf{F}}}\chi \quad (33)$$

for

$$X = \frac{2\{1+\omega\}}{2\omega+3}. \quad (34)$$

⁹Were tilt terms not removable, as these contain spatial derivatives of $\dot{\mathbf{I}}$, they would compromise the algebraicity of the elimination of $\dot{\mathbf{I}}$ from the cyclic equation that arises from variation with respect to $\dot{\mathbf{I}}$. Of course, the algebraic equation might have no roots or only physically unacceptable (e.g. non-real) roots, in which cases the theory should be discarded. It might also not be explicitly soluble – that is what I mean by ‘in principle’.

¹⁰Under this field redefinition, it is then a scalar field minimally coupled to gravity, which is clearly included among the 3-space approach castable cases listed in Sec 1.5.

Thus it is equivalence principle violating as it contains metric–matter kinetic cross terms. Nevertheless it can be cast into 3-space approach form [33] (but was missed in [4] through the ansatz there not including metric–matter kinetic cross-terms).

However, this example suffers the observational weakness that its parameter ω is fixed, expected on grounds of theoretical naturality to be of order unity and yet is bounded by the Cassini data to be above 20000 [41]. This weakness can be removed by showing that the more general scalar–tensor theory with spacetime action¹¹

$$I_{\text{STT}}[g_{AB}, \chi] = \int d^4x \sqrt{|g|} e^{-\chi/2} \{ \mathcal{R} - \omega(\chi) \partial_A \chi \partial^A \chi + \mathbf{U}(\chi) \} \quad (35)$$

which one can likewise cast as a 3-space approach theory by performing the split with respect to a family of spatial hypersurfaces using instant–grid variables and then eliminating $\dot{\mathbf{I}}$ and writing $\mathcal{L}_{\dot{\mathbf{F}}}$ for $\delta_{\dot{\mathbf{F}}}$ in the usual fashion. That this can be cast in 3-space approach form is clear because adding a potential and replacing ω with $\omega(\chi)$ do not affect the split of the spacetime tensorial objects in the action or the form that the Routhian reduction that eliminates $\dot{\mathbf{I}}$ are to take. While, this no longer suffers from the observational weakness because now ω varies and there is evidence that it tends dynamically to a large value in the late universe (toward the general relativity value of $+\infty$) [43].

3.4 Vector–tensor theories

An example of 3-space approach theory [15] in which the second kind of equivalence principle violation occurs can be found among the the vector–tensor theories considered in e.g. [44]. This class of theories has the spacetime form:

$$I_{\text{VTT}}[g_{AB}, A_A] = \int d\lambda \int d^3x \alpha \{ \mathcal{R} + \nu \{ \nabla_A A^A \nabla_B A^B + m^2 A^2 \} \} \quad (36)$$

Now the split form of action (36) is (by the above derivative formulae and then using the field redefinition

$$\dot{\mathbf{I}}A^a = \dot{v} \quad (37)$$

to remove ‘tilts’ and also setting A_{\perp} to be some ϕ):

$$I_{\text{VTT}}^{\text{ADM}}[h_{ab}, \dot{h}_{ab}, v_i, \dot{v}_i, \phi, \dot{\phi}, \dot{\mathbf{F}}_i, \dot{\mathbf{I}}] = \int \int d\lambda \dot{\mathbf{I}} \sqrt{\dot{h}} d^3x \left\{ \frac{\mathbf{T}_{\text{GR}}^{\mathbf{A}}[h_{ab}, \dot{h}_{ab}, \dot{\mathbf{F}}_i]}{4\dot{\mathbf{I}}^2} + \frac{\nu}{\dot{\mathbf{I}}^2} \left\{ \left\{ D_a \{ \dot{v}^a \} + \frac{\phi}{2} h^{ij} \delta_{\dot{\mathbf{F}}} h_{ij} + \delta_{\dot{\mathbf{F}}} \phi \right\}^2 + m^2 \dot{v}^2 \right\} + R - \nu m^2 \phi^2 \right\}. \quad (38)$$

Then a Routhian reduction of the same form as that mentioned in Sec 1.2 is possible, giving

$$I_{\text{VTT}}^{\mathbf{A}'}[h_{ab}, \dot{h}_{ab}, v_i, \dot{v}_i, \phi, \dot{\phi}, \dot{\mathbf{F}}_i] = \int \int d\lambda d^3x \sqrt{\dot{h}} \sqrt{ \{ R - \nu m^2 \phi^2 \} \left\{ \mathbf{T}_{\text{GR}}^{\mathbf{A}'} + 4\nu \left\{ \left\{ D_a \{ \dot{v}^a \} + \frac{\phi}{2} h^{ij} \delta_{\dot{\mathbf{F}}} h_{ij} + \delta_{\dot{\mathbf{F}}} \phi \right\}^2 + m^2 \dot{v}^2 \right\} \right\} }. \quad (39)$$

[The equations encoded by this action happen to be weakly unaffected by whether \dot{v}^a is replaced by $\delta_{\dot{\mathbf{F}}} v^a$.]

¹¹This is not the most general scalar-tensor theory (see e.g. [42] and references therein). E.g. one could replace $e^{-\chi/2}$ by an arbitrary function of χ , or furthermore extend the theory to have more than 1 scalar. However, the example in this paper is general enough to illustrate the point in question.

Thus if one starts with 3-space approach principles, and using the arbitrary 3-diffeomorphism frame symbol $\&_{\dot{F}}$ in place of the hypersurface derivative symbol $\delta_{\dot{F}}$, one obtains the 3-space approach action

$$\begin{aligned} & \mathfrak{I}_{\sqrt{\text{TT}}}^{\text{TSA}}[h_{ab}, \dot{h}_{ab}, v_i, \dot{v}_i, \phi, \dot{\phi}, \dot{F}_i] = \\ & \int \int d\lambda d^3x \sqrt{\hbar} \sqrt{\{R - \nu m^2 \phi^2\} \left\{ \mathfrak{T}_{\text{GR}}[h_{ab}, \dot{h}_{ab}, \dot{F}_i] + 4\nu \left\{ \left\{ D_a \{ \&_{\dot{F}} v^a \} + \frac{\phi}{2} h^{ij} \&_{\dot{F}} h_{ij} + \&_{\dot{F}} \phi \right\}^2 + m^2 \{ \&_{\dot{F}} v \}^2 \right\} \right\}}. \end{aligned} \quad (40)$$

Thus one has a consistent (by reverse of above working and the original spacetime formulation being consistent) and nontrivial equivalence principle violating theory for geometry, a scalar and a 1-form. It should be noted that [4] missed this not on relational grounds but on simplicity grounds: the theory has a kinetic term that is not ultralocal, has metric–matter cross-terms and field dependence.

This was missed in [4] through it having metric–matter kinetic cross-terms, matter field dependence in the kinetic metric and a mixture of 1-form and scalar modes from the 3-space perspective.

Many theories of this type have a number of undesirable features, such as classical and quantum instabilities [44, 45], non-positivness of total energy [47] and formation of shocks beyond which the evolution cannot be extended [45]. [15] speculated that some axiom that avoids such pathologies could be used to bring down this class of counterexample.

4 A new example of 3-space approach theory that is all of equivalence principle violating, special relativity violating and less pathological

4.1 Unit vector–tensor theories (Einstein–Aether theories)

[46, 47, 48] consider a general Einstein–Aether action of the form

$$\mathfrak{I}_{\text{EAT}}[g_{AB}, u_A] = \int d^4x \sqrt{-g} \{ \mathcal{R} + E_1 \{ \nabla_A u_B \} \nabla^A u^B + E_2 \{ \nabla_A u^A \}^2 + E_3 \{ \nabla_A u^B \} \nabla_B u^A + E_4 u^A u^B \{ \nabla_A u_C \} \nabla_B u^C + \lambda \{ u_A u^A - 1 \} \}. \quad (41)$$

As compared to the general theories considered by Isenberg and Nester, this permits 1 further derivative term (though I do not make use of it in my specific examples), and furthermore interprets what was the mass now as a Lagrange multiplier and adds the multiplier again as an extra potential piece. [The Lagrange multiplier is there to implement the unit-field constraint.] At least some of these unit-field theories are less pathological [47].

These theories are in general equivalence principle violators, the exception being if all of

$$E_1 + E_3 = 0 \quad (\text{Maxwellian combination}), \quad (42)$$

$$E_2 = 0 = E_4 \quad (43)$$

hold. These theories are also in general special relativity violating, for they contain [46, 47] spin-2 fields propagating at squared speeds

$$c_2 = \frac{1}{1 - \{E_1 + E_3\}}, \quad (44)$$

spin-1 fields propagating at speed

$$c_1 = \frac{E_1 - E_1^2/2 + E_3^2/2}{\{E_1 + E_4\} \{1 - \{E_1 + E_3\}\}} \quad (45)$$

and spin-0 fields propagating at speed

$$c_0 = \frac{\{E_1 + E_2 + E_3\}\{2 - \{E_1 + E_4\}\}}{\{E_1 + E_4\}\{1 - \{E_1 + E_3\}\}\{2 + E_1 + E_3 + 3E_2\}} . \quad (46)$$

These are fairly extensively finite and with at least one distinct from the speed of light $c = 1$ in these units. This is the case unless all of

$$E_1 + E_3 = 0 , \quad (47)$$

$$E_4 = 0 , \quad (48)$$

and

$$E_1^{-1} - E_2^{-1} = 2 \quad (49)$$

hold.

These squared speeds are also capable of going negative, corresponding to undesirable exponential-type instabilities. Positive linearized energy density requires

$$\{2E_1 - E_1^2 + E_3^2\}/\{1 - E_1 - E_3\} > 0 \quad (50)$$

(vector mode contribution) and

$$\{E_1 + E_4\}\{2 - E_1 - E_4\} > 0 \quad (51)$$

(trace mode contribution). One would also like the kinetic energy contributions to have the usual sign for matter kinetic terms.

4.2 Einstein–Aether theories that are castable in 3-space approach form

To build a suitable 3-space approach example that is equivalence principle violating, special relativity violating in the sense of having 2 different finite fundamental propagation speeds and not subject to the above three pathologies, proceed as follows.

Consider first the theory with E_2 alone nonzero. Compared to the previous section's theory, the only difference is to the potential (which is trivial to split spacetime framework decompose), so the previous section's working will straightforwardly extend to the Einstein–aether theory case that is analogous to the above specially-chosen case. But for E_2 theory $c_2 = 1 (= c)$ and the other two are not finite, so this does not constitute a special relativity violation of the type I am seeking.

But consider then furthermore including a Maxwell-type combination ($E_1 = -E_3 \neq 0$) in the action; as this has a very simple split spacetime framework decomposition, this addition does not ruin the algebraicity of the Routhian reduction. So the theory I choose to work with is identified in the spacetime picture as

$$\mathcal{I}_{\text{EAT}}[g_{AB}, u_A] = \int d^4x \sqrt{-g} \{ \mathcal{R} + E_1 \{ \nabla_A u_B \} \nabla^A u^B + E_2 \{ \nabla_A u^A \}^2 - E_1 \{ \nabla_A u^B \} \nabla_B u^A + \lambda \{ u_A u^A - 1 \} \} . \quad (52)$$

split spacetime framework splitting this, adhering to the redefinition (37), using symmetry-antisymmetry cancellations on the new quadratic tilt terms and integration by parts on the new linear tilt terms, one indeed passes to a homogeneous quadratic action to which the usual Routhian reduction move can be carried out. The resulting action may, moreover be interpreted as (rewriting $\delta_{\dot{\mathcal{F}}}$ as $\&\dot{\mathcal{F}}$ and adopting this action as one's new starting-point) a 3-space approach theory that follows from the configurational and temporal relationalism principles:

$$\mathcal{I}_{\text{EAT}}^{\text{TSA}}[h_{ab}, \dot{h}_{ab}, \dot{\mathcal{F}}_a, \dot{v}_a] = \int d\lambda \int d^3x \sqrt{h} \sqrt{\mathcal{T}\mathcal{U}} \quad (53)$$

for

$$\mathbf{U} = R - \mu\{\phi^2 + 1\} \quad (54)$$

and

$$\mathbb{T} = \mathbb{T}_{\text{GR}}^{\text{BFO-A}} + \mathbb{T}_{\mathbf{v}}^{\text{BFO-A}}(\nu \rightarrow E_2, m^2 \rightarrow \mu) + \mathbb{T}_{\mathbf{v}}^{\text{BFO-A}'}, \quad (55)$$

where

$$\mathbb{T}_{\mathbf{v}}^{\text{BFO-A}'} = E_1 \{ \{ h^{ac} h^{bd} - h^{ad} h^{bc} \} \partial_a \{ \&_{\mathbf{F}} v_b \} \partial_c \&_{\mathbf{F}} v_d + 2D^b \{ \&_{\mathbf{F}} v^a \{ \partial_b \&_{\mathbf{F}} v_a - \partial_a \&_{\mathbf{F}} v_b \} \} + \{ \&_{\mathbf{F}} \phi \}^2 \} . \quad (56)$$

Compared to the original example I gave, this is more general in having E_2 and less general in being a unit vector field.

Now, indeed, there is in general more than 1 fundamental propagation speed, as

$$c_2 = c_1 = 1 \quad (= c), \quad c_0 = \frac{E_2 \{ 2 - E_1 \}}{E_1 \{ 2 + 3E_2 \}} . \quad (57)$$

I.e. this example contains 1) a non-generic case

$$E_1^{-1} - E_2^{-1} = 2 \quad (58)$$

which is *not* a counterexample to violation of the special relativity lightcone (this is a subcase of the above non-special relativity violating example). 2) The general case

$$E_1^{-1} - E_2^{-1} \neq 2 \quad (59)$$

for which

$$c_0 \neq 1 = c_1 = c_2 = c . \quad (60)$$

So there is a 1-parameter family (bar a single parameter value) of 3-space approach complying special relativity violating theories: there are scalar modes whose propagation speed in vacuo is different from the speed of light, so these have a null cone structure that is *not* shared with the other fields in this theory.

A fair portion of the above example's parameter space is able to comply with positive linearized energy: that for which

$$0 < E_1 < 2 . \quad (61)$$

For $E_2 > 0$ or $< -2/3$, this remaining region complies with the stability criteria $c_0^2, c_1^2, c_2^2 > 0$ since these reduce to the trivial $1 > 0$ (twice) and

$$\frac{E_2 \{ 2 - E_1 \}}{E_1 \{ 2 + 3E_2 \}} > 0 , \quad (62)$$

however the latter subregion should be discarded to ensure that the kinetic term is of the right characteristic sign for a matter contribution, thus leaving one with the 'region of non-pathology'

$$0 < E_1 < 2 , \quad E_2 > 0 \quad (63)$$

for the theory's coupling constants.

This region is split into two pieces by the curve of special relativity-compliance [i.e. of universal luminal fundamental speed (49)]; the subregion above this curve has the scalar modes propagate superluminally and the subregion below this curve has them propagate subluminally.

5 Conclusion

The 3-space approach is based on temporal and configurational relational principles. General relativity in geometrodynamical form can be derived as one consistent alternative that follows from these premises when applied to a theory for which the 3-metrics on a fixed spatial topology are redundant dynamical objects under the associated 3-diffeomorphisms. A sufficient set of fundamental matter fields to describe nature can be adjoined to this scheme. It was furthermore claimed that

- 1) working with matter fields alongside spatial 3-metrics *picks out* electromagnetism (and Yang–Mills theory) coupled to general relativity as the only consistent theories of one (and K interacting) 1-forms.
- 2) The equivalence principle is emergent.
- 3) The universal null cone of special relativity is locally recovered.

These were always subject to simplicity assumptions as well as the relational postulates. Claim 1) should be weakened, at least on basis of current workings. This is not only because lifting unrelational simplicities that were identified as such at the time of doing the calculation has been shown to destroy the result, but also because of two further tacit simplicities assumed in the proof, one of which is unmotivated and the other of which is unduly restrictive from a theoretical perspective. Without these the exhaustion goes more slowly and one then has to work case by case rather than once and for all with an arbitrary gauge group.

Also, examples including some new to this paper show that it is necessary for the relational postulates to be supplemented by non-relational simplicity assumptions in order for 2) and 3) to hold. As these necessary simplicity assumptions include what Hojman, Kuchař and Teitelboim identify as the geometrodynamical equivalence principle, the 3-space approach’s claim of *deriving* the equivalence principle loses its credibility. Furthermore, from the split spacetime framework perspective, the geometrodynamical equivalence principle and statements equivalent to the relational postulates come as a neat package involving the three types of universal kinematics: hypersurface derivatives, the absence of tilts and the absence of derivative couplings, while taking the 3-space approach and the geometrodynamical equivalence principle as one’s principles is more heterogeneous.¹² That said, the equivalence principle is separate from other postulates in Einstein-type spacetime approach, so one is doing no worse than what one does in taking the geometrodynamical equivalence principle alongside the relational postulates to be the heart of the axiomatization of general relativity. That reflects the primality of the equivalence principle as regards axiomatizations of general relativity – so far as the author (or Brown [49]) are aware, no derivations of the equivalence principle from more basic postulates are known (which is what merited my concentrated effort to bring down [4]’s conjecture otherwise).

‘Simple’ (in the sense of Sec 1.5) matter fields coupled to dynamical 3-metrics builds in the equivalence principle; one then encounters (an extension of) the fork Einstein encountered in setting up special relativity as the “roots of” an explicit equation arising from the Hamiltonian-type constraint by the Dirac procedure. These correspond to Lorentzian relativity (single finite physical propagation speed), Galilean relativity (infinite propagation speed), and Carrollian relativity (zero propagation speed). But if the associated simplicities are dropped, this article’s example shows that equivalence principle violation is possible including in otherwise relatively non-pathological situations, and more than 1 finite fundamental propagation speed can occur – consistency other than by the above fork becomes allowed.

A new issue to investigate – I dare not call it a conjecture – is whether each of the local recovery of special relativity and of gauge theory can be shown to follow from the equivalence principle free of the 3-space approach or even geometrodynamical formalism. Here, ‘shown to follow’ might mean that they are among the natural structures to emerge, and perhaps a further axiom or a collection of observations or demands from local quantum field theory could remove some (or all) of the structures that co-emerge with them. This does not look to be restricted to a specific geometrodynamical formulation or even to geometrodynamics, but would rather be a stronger formalism-independent result of general relativity.

¹²The 3-space approach assumes less structure than (split) spacetime approaches. That makes it ‘more interesting’ but also harder to work with as there being less structure makes proving theorems harder in the 3-space approach than e.g. in Hojman–Kuchař–Teitelboim’s approach that presupposes spacetime. E.g., their use of induction proofs specifically rely on additional spacetime structure.

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