DESINGULARIZATION OF 3D STEADY EULER EQUATIONS WITH HELICAL SYMMETRY

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ABSTRACT. In this paper, we study desingularization of steady solutions of 3D incompressible Euler equation with helical symmetry in a general helical domain. We construct a family of steady Euler flows with helical symmetry, such that the associated vorticities tend asymptotically to a helical vortex filament. The solutions are obtained by solving a semilinear elliptic problem in divergence form with a parameter. By using the streamfunction method, we show the existence and asymptotic behavior of ground state solutions concentrating near a single point as the parameter $\varepsilon \to 0$. Qualitative properties of those solutions are also discussed.

Keywords: Desingularization; Steady Euler equations; Helical symmetry; Variational method.

1. Introduction and main results

1.1. **Introduction.** The movement of incompressible Euler flow confined in a three-dimensional domain D without external force is governed by the following system

$$\begin{cases}
\partial_t \mathbf{v} + (\mathbf{v} \cdot \nabla) \mathbf{v} = -\nabla P, & D \times (0, T), \\
\nabla \cdot \mathbf{v} = 0, & D \times (0, T), \\
\mathbf{v} \cdot \mathbf{n} = v_n, & \partial D \times (0, T), \\
\mathbf{v}(x, 0) = \mathbf{v}_0(x), & D,
\end{cases}$$
(1.1)

where $D \subseteq \mathbb{R}^3$ is a domain with C^{∞} boundary, $\mathbf{v} = (v_1, v_2, v_3)$ is the velocity field, P is the scalar pressure, \mathbf{n} is the outward unit normal of ∂D and \mathbf{v}_0 is the initial velocity. v_n is a function defined on ∂D satisfying compatibility condition

$$\int_{\partial D} v_n d\sigma = 0,$$

where σ is the area unit on ∂D . The third equation of (1.1) means that the net flux of velocity across the boundary is zero. When $v_n \equiv 0$, it is the impermeable boundary condition, which means that the normal component of velocity on the boundary is zero.

The vorticity vector field associated with \mathbf{v} is $\mathbf{w} = (w_1, w_2, w_3) = curl\mathbf{v} = \nabla \times \mathbf{v}$, which describes the rotation of the fluid. Then \mathbf{w} satisfies the vorticity equations

$$\begin{cases}
\partial_{t}\mathbf{w} + (\mathbf{v} \cdot \nabla)\mathbf{w} = (\mathbf{w} \cdot \nabla)\mathbf{v}, & D \times (0, T), \\
\nabla \cdot \mathbf{v} = 0, & D \times (0, T), \\
\mathbf{v} \cdot \mathbf{n} = v_{n}, & \partial D \times (0, T), \\
\mathbf{w}(x, 0) = \nabla \times \mathbf{v}_{0}(x), & D.
\end{cases}$$
(1.2)

For background of the 3D incompressible Euler equation, see the classical literature [21, 22]. In this paper, we are devoted to Euler equations (1.1) with helical symmetry. Let us first define helical symmetric solutions and simplify the vorticity equations (1.2), see [9, 12, 13]. Let k > 0. Define a one-parameter group $\mathcal{G}_k = \{H_\rho : \mathbb{R}^3 \to \mathbb{R}^3\}$, where

$$H_{\rho}(x_1, x_2, x_3)^t = (x_1 \cos \rho + x_2 \sin \rho, -x_1 \sin \rho + x_2 \cos \rho, x_3 + k\rho)^t.$$

Here A^t is the transposition of a matrix A. So H_{ρ} is a superposition of a rotation in x_1Ox_2

plane and a translation in
$$x_3$$
 axis. Let $R_{\rho} = \begin{pmatrix} \cos \rho & \sin \rho & 0 \\ -\sin \rho & \cos \rho & 0 \\ 0 & 0 & 1 \end{pmatrix}$ be the rotation with

respect to x_3 -axis. Then $H_{\rho}(x) = R_{\rho}(x) + k\rho(0,0,1)$. From a geometric point of view, $2k\pi$ corresponds to the pitch of helices.

Define a vector field

$$\overrightarrow{\zeta} = (x_2, -x_1, k)^t.$$

Then $\overrightarrow{\zeta}$ is the field of tangents of symmetry lines of \mathcal{G}_k .

To show what the helical solutions are, we first define helical domains. A domain $D \in \mathbb{R}^3$ is called a *helical domain*, if $H_{\rho}(D) = D$ for any ρ . So D is invariant under the group \mathcal{G}_k . Let $\Omega = D \cap \{x \mid x_3 = 0\}$ be the section of D over x_1Ox_2 plane. Then D can be generated by Ω by letting $D = \bigcup_{\rho \in \mathbb{R}} H_{\rho}(\Omega)$. Throughout this paper, we always assume that Ω is a simply-connected bounded domain with C^{∞} boundary and D is a helical domain generated by Ω .

Now we give the definition of helical functions and vector fields. A scalar function h is called a *helical* function, if

$$h(H_{\rho}(x)) = h(x) \tag{1.3}$$

for any $\rho \in \mathbb{R}, x \in D$. By direct computations it is easy to see that a C^1 function h is helical if and only if

$$\overrightarrow{\zeta} \cdot \nabla h = 0.$$

A vector field $\mathbf{h} = (h_1, h_2, h_3)$ is called a *helical* field, if

$$\mathbf{h}(H_{\rho}(x)) = R_{\rho}\mathbf{h}(x) \tag{1.4}$$

for any $\rho \in \mathbb{R}, x \in D$. Direct computation shows that a C^1 vector field **h** is helical if and only if

$$\overrightarrow{\zeta} \cdot \nabla \mathbf{h} = \mathcal{R} \mathbf{h},$$

where $\mathcal{R} = \begin{pmatrix} 0 & 1 & 0 \\ -1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}$ (see [13]). Helical solutions of (1.1) are then defined as follows.

Definition 1.1. A function pair (\mathbf{v}, P) is called a *helical* solution pair of (1.1), if (\mathbf{v}, P) satisfies (1.1) and both vector field \mathbf{v} and scalar function P are helical.

Throughout this paper, helical solutions also need to satisfy the *orthogonality condition*:

$$\mathbf{v} \cdot \overrightarrow{\zeta} = 0, \tag{1.5}$$

that is, the velocity field and $\overrightarrow{\zeta}$ are orthogonal.

Under the condition (1.5), one can check that the vorticity field **w** satisfies (see [13])

$$\mathbf{w} = \frac{w \to}{k} \zeta, \tag{1.6}$$

where $w := w_3 = \partial_{x_1} v_2 - \partial_{x_2} v_1$, the third component of vorticity field **w**, is a helical function. Moreover, the first equation of the vorticity equations (1.2) is equivalent to

$$\partial_t \mathbf{w} + (\mathbf{v} \cdot \nabla) \mathbf{w} + \frac{1}{k} w \mathcal{R} \mathbf{v} = 0.$$

As a consequence, w satisfies

$$\partial_t w + (\mathbf{v} \cdot \nabla) w = 0. \tag{1.7}$$

From (1.7) we deduce that, w satisfies a transport equation, which is very similar to the case of 2D Euler equations (see Yudovich [26]). Moreover, for a solution w of (1.7), the vorticity field \mathbf{w} is determined by (1.6).

We now introduce a *stream function* and reduce the system (1.2) to a 2D problem. Since \mathbf{v} is a helical vector field, we have $\overrightarrow{\zeta} \cdot \nabla \mathbf{v} = \mathcal{R} \mathbf{v}$, which implies that

$$x_2 \partial_{x_1} v_3 - x_1 \partial_{x_2} v_3 + k \partial_{x_3} v_3 = 0. (1.8)$$

The orthogonal condition shows that

$$x_2v_1 - x_1v_2 + kv_3 = 0. (1.9)$$

It follows from the incompressible condition, (1.8) and (1.9) that

$$0 = \partial_{x_1} v_1 + \partial_{x_2} v_2 + \partial_{x_3} v_3 = \partial_{x_1} v_1 + \partial_{x_2} v_2 - \frac{x_2}{k} \partial_{x_1} v_3 + \frac{x_1}{k} \partial_{x_2} v_3$$

$$= \partial_{x_1} v_1 + \partial_{x_2} v_2 - \frac{x_2}{k^2} \partial_{x_1} (-x_2 v_1 + x_1 v_2) + \frac{x_1}{k^2} \partial_{x_2} (-x_2 v_1 + x_1 v_2)$$

$$= \frac{1}{k^2} \partial_{x_1} [(k^2 + x_2^2) v_1 - x_1 x_2 v_2] + \frac{1}{k^2} \partial_{x_2} [(k^2 + x_1^2) v_2 - x_1 x_2 v_1].$$

Since Ω is simply-connected, we can define a stream function $\varphi:\Omega\to\mathbb{R}$ such that $\partial_{x_2}\varphi=\frac{1}{k^2}[(k^2+x_2^2)v_1-x_1x_2v_2], \partial_{x_1}\varphi=-\frac{1}{k^2}[(k^2+x_1^2)v_2-x_1x_2v_1]$, that is,

$$\begin{pmatrix} \partial_{x_1} \varphi \\ \partial_{x_2} \varphi \end{pmatrix} = -\frac{1}{k^2} \begin{pmatrix} -x_1 x_2 & k^2 + x_1^2 \\ -(k^2 + x_2^2) & x_1 x_2 \end{pmatrix} \begin{pmatrix} v_1 \\ v_2 \end{pmatrix},$$

or equivalently,

$$\begin{pmatrix} v_1 \\ v_2 \end{pmatrix} = -\frac{1}{k^2 + x_1^2 + x_2^2} \begin{pmatrix} x_1 x_2 & -k^2 - x_1^2 \\ k^2 + x_2^2 & -x_1 x_2 \end{pmatrix} \begin{pmatrix} \partial_{x_1} \varphi \\ \partial_{x_2} \varphi \end{pmatrix}. \tag{1.10}$$

By the definition of w and (1.10), we get

$$w = \partial_{x_{1}} v_{2} - \partial_{x_{2}} v_{1} = (-\partial_{x_{2}}, \partial_{x_{1}}) \begin{pmatrix} v_{1} \\ v_{2} \end{pmatrix}$$

$$= (-\partial_{x_{2}}, \partial_{x_{1}}) \left(-\frac{1}{k^{2} + x_{1}^{2} + x_{2}^{2}} \begin{pmatrix} x_{1} x_{2} & -k^{2} - x_{1}^{2} \\ k^{2} + x_{2}^{2} & -x_{1} x_{2} \end{pmatrix} \begin{pmatrix} \partial_{x_{1}} \varphi \\ \partial_{x_{2}} \varphi \end{pmatrix} \right)$$

$$= -(\partial_{x_{1}}, \partial_{x_{2}}) \left(\frac{1}{k^{2} + x_{1}^{2} + x_{2}^{2}} \begin{pmatrix} k^{2} + x_{2}^{2} & -x_{1} x_{2} \\ -x_{1} x_{2} & k^{2} + x_{1}^{2} \end{pmatrix} \begin{pmatrix} \partial_{x_{1}} \varphi \\ \partial_{x_{2}} \varphi \end{pmatrix} \right)$$

$$= \mathcal{L}_{H} \varphi,$$

$$(1.11)$$

where $\mathcal{L}_H \varphi = -\text{div}(K_H(x_1, x_2)\nabla\varphi)$ is a second order elliptic operator of divergence type with the coefficient matrix

$$K_H(x_1, x_2) = \frac{1}{k^2 + x_1^2 + x_2^2} \begin{pmatrix} k^2 + x_2^2 & -x_1 x_2 \\ -x_1 x_2 & k^2 + x_1^2 \end{pmatrix}.$$
(1.12)

Clearly from the definition of the matrix K_H , K_H is a positive definite matrix satisfying

- $(\mathcal{K}1)$. K_H is smooth, i.e., $(K_H(\cdot))_{ij} \in C^{\infty}(\overline{\Omega})$ for i, j = 1, 2.
- ($\mathcal{K}2$). \mathcal{L}_H is strictly elliptic. Indeed, two eigenvalues of K_H are $\lambda_1 = 1, \lambda_2 = \frac{k^2}{k^2 + |x|^2}$. So one has,

$$\frac{k^2}{k^2 + |x|^2} |\zeta|^2 \le (K_H(x)\zeta|\zeta) \le |\zeta|^2, \quad \forall \ x \in \Omega, \ \zeta \in \mathbb{R}^2.$$

From (1.7), (1.9) and (1.10), one has

$$0 = \partial_{t}w + v_{1}\partial_{x_{1}}w + v_{2}\partial_{x_{2}}w + v_{3}\partial_{x_{3}}w$$

$$= \partial_{t}w + v_{1}\partial_{x_{1}}w + v_{2}\partial_{x_{2}}w + \frac{1}{k}(-x_{2}v_{1} + x_{1}v_{2}) \cdot \frac{1}{k}(-x_{2}\partial_{x_{1}}w + x_{1}\partial_{x_{2}}w)$$

$$= \partial_{t}w + \frac{1}{k^{2}}(v_{1}, v_{2}) \begin{pmatrix} k^{2} + x_{2}^{2} & -x_{1}x_{2} \\ -x_{1}x_{2} & k^{2} + x_{1}^{2} \end{pmatrix} \begin{pmatrix} \partial_{x_{1}}w \\ \partial_{x_{2}}w \end{pmatrix}$$

$$= \partial_{t}w - \frac{1}{k^{2}(k^{2} + x_{1}^{2} + x_{2}^{2})} (\partial_{x_{1}}\varphi, \partial_{x_{2}}\varphi) \begin{pmatrix} x_{1}x_{2} & k^{2} + x_{2}^{2} \\ -k^{2} - x_{1}^{2} & -x_{1}x_{2} \end{pmatrix} \begin{pmatrix} k^{2} + x_{2}^{2} & -x_{1}x_{2} \\ -x_{1}x_{2} & k^{2} + x_{1}^{2} \end{pmatrix} \begin{pmatrix} \partial_{x_{1}}w \\ \partial_{x_{2}}w \end{pmatrix}$$

$$= \partial_{t}w - (\partial_{x_{1}}\varphi, \partial_{x_{2}}\varphi) \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \begin{pmatrix} \partial_{x_{1}}w \\ \partial_{x_{2}}w \end{pmatrix}$$

$$= \partial_{t}w + \partial_{x_{2}}\varphi\partial_{x_{1}}w - \partial_{x_{1}}\varphi\partial_{x_{2}}w$$

$$= \partial_{t}w + \nabla w \cdot \nabla^{\perp}\varphi,$$

$$(1.13)$$

where \perp denotes the clockwise rotation through $\pi/2$.

As for the boundary condition of φ , from the fact that \mathbf{v} is a helical vector field and the domain D is helical, v_n is a helical function defined on ∂D . So v_n can be generated by $v_n|_{\partial\Omega}$. Moreover, it follows from $\mathbf{v} \cdot \mathbf{n} = v_n$ on ∂D that (see (2.66), [13])

$$\begin{aligned} v_n|_{\partial\Omega} &= \mathbf{v} \cdot \mathbf{n}|_{\partial\Omega} \\ &= v_1 n_1 + v_2 n_2 + \left(-\frac{x_2}{k} v_1 + \frac{x_1}{k} v_2 \right) \left(-\frac{x_2}{k} n_1 + \frac{x_1}{k} n_2 \right) \\ &= \frac{1}{k^2} (v_1, v_2) \begin{pmatrix} k^2 + x_2^2 & -x_1 x_2 \\ -x_1 x_2 & k^2 + x_1^2 \end{pmatrix} \begin{pmatrix} n_1 \\ n_2 \end{pmatrix}. \end{aligned}$$

Combining this with (1.10), one has

$$v_{n}|_{\partial\Omega} = -\frac{1}{k^{2}(k^{2} + x_{1}^{2} + x_{2}^{2})} (\partial_{x_{1}}\varphi, \partial_{x_{2}}\varphi) \begin{pmatrix} x_{1}x_{2} & k^{2} + x_{2}^{2} \\ -k^{2} - x_{1}^{2} & -x_{1}x_{2} \end{pmatrix} \begin{pmatrix} k^{2} + x_{2}^{2} & -x_{1}x_{2} \\ -x_{1}x_{2} & k^{2} + x_{1}^{2} \end{pmatrix} \begin{pmatrix} n_{1} \\ n_{2} \end{pmatrix}$$

$$= -(\partial_{x_{1}}\varphi, \partial_{x_{2}}\varphi) \begin{pmatrix} 0 & 1 \\ -1 & 0 \end{pmatrix} \begin{pmatrix} n_{1} \\ n_{2} \end{pmatrix}$$

$$= \nabla^{\perp}\varphi \cdot \nu. \tag{1.14}$$

Here $\nu = (n_1, n_2)$ is the two-dimensional vector of the first two component of \mathbf{n} on $\partial\Omega$. Note that ν is an outward normal of $\partial\Omega$.

Thus, the 2D vorticity equations of 3D Euler equations with helical symmetry is as follows

$$\begin{cases} \partial_t w + \nabla w \cdot \nabla^{\perp} \varphi = 0, \\ w = \mathcal{L}_H \varphi, \\ \nabla^{\perp} \varphi \cdot \nu|_{\partial \Omega} = v_n|_{\partial \Omega}. \end{cases}$$
 (1.15)

Indeed for a solution pair (w, φ) of the 2D vorticity equation (1.15), one can recover the helical velocity field \mathbf{v} and vorticity field \mathbf{w} of 3D Euler equations (1.1) by using (1.10), (1.9), (1.4), (1.3) and (1.6).

When considering the case $v_n \equiv 0$, which corresponds to the impermeable boundary condition, (1.14) implies that φ is a constant on $\partial\Omega$. Thus one can choose φ such that $\varphi \equiv 0$ on $\partial\Omega$. The associated vorticity equations then become

$$\begin{cases} \partial_t w + \nabla w \cdot \nabla^{\perp} \varphi = 0, \\ w = \mathcal{L}_H \varphi, \\ \varphi|_{\partial \Omega} = 0. \end{cases}$$
 (1.16)

The research of 3D Euler equations with helical symmetry has received much attention in recent years. [13] first proved the global well-posedness of $L^1 \cap L^{\infty}$ weak solutions of the Euler equation with helical symmetry without vorticity stretching (see (1.16)). Note that this result corresponds to the classical Yudovich's result [26], since the structure of vorticity equation (1.16) is similar to that of 2D Euler flows. [10] considered travelling-rotating invariant Euler flows with helical symmetry concentrating near a single helical filament in the whole space \mathbb{R}^3 . As for the steady solution of 3D Euler equations with helical symmetry,

nonlinear stability for stationary smooth Euler flows with helical symmetry is considered in [2] by using the direct method of Lyapunov. More results of the existence and regularity of Euler equation with helical symmetry can be found in [1, 4, 12, 19] for instance.

In this paper we are interested in vortex desingularization problem of steady 3D Euler equations with helical symmetry. Here the vortex desingularization problem means that we want to construct a family of "true" solutions of Euler equations, such that the corresponding vorticity has a small cross-section and concentrates near a vortex filament. The research of this problem can be traced back to Helmholtz [15], who first studied the motion of the travelling vortex rings whose vorticities are supported in toroidal regions with a small cross-section. Then, many articles considered the problem. As for the vortex being a tube with a small cross-section whose centerline is a straight line and a circle, which can be reduced to 2D Euler equation and 3D axisymmetric Euler equation respectively, results can be found in [3, 5, 6, 9, 11, 14, 24] and reference therein. For the case of Euler equations with helical symmetry, results seem to be few. Dávila et al. [10] constructed rotational-invariant Euler flows with helical symmetry in the whole space by considering

$$-\operatorname{div}(K_H(x)\nabla u) = f_{\varepsilon}(u - \alpha|\ln \varepsilon|\frac{|x|^2}{2}) \text{ in } \mathbb{R}^2,$$

where $f_{\varepsilon}(t) = \varepsilon^2 e^t$. Using the Lyapunov-Schmidt reduction method, the authors proved that solutions will concentrate near a helix in the distributional sense, which satisfies the vortex filament conjecture, see [17, 18]. Desingularization of rotational-invariant helical solutions in helical domains with bounded cross section was proved in [8]. However for the problem of desingularization of steady Euler equations with helical symmetry, few results give us a positive answer.

Our goal in this article is to solve vortex desingularization problem of steady Euler equations with helical symmetry in general helical domains. We will construct steady solutions of vorticity equations (1.15) and (1.16), such that the associated vorticities have small cross-sections and concentrate near a single point as parameter changes. Accordingly, the vorticity field will concentrate near a one-dimensional helical filament. Note that both the case of $v_n \equiv 0$ and that of $v_n \not\equiv 0$ are considered. To get these results, we solve solutions of a semilinear elliptic equation in divergence form (see (2.1)). By studying the associated variational structure and using stream function method, we get the existence and limiting behavior of ground states of these equations.

It should be noted that, Euler equations with helical symmetry can be regarded as the general case of 2D and 3D axisymmetric Euler equations. The cases $k \to +\infty$ and k = 0 correspond to the 2D Euler equations and 3D axisymmetric Euler equations, respectively. The case $k \in (0, +\infty)$ is considered in this paper. In contrast to the 2D and 3D axisymmetric problem, the associated operator \mathcal{L}_H in vorticity equations (1.15) is a general elliptic operator in divergence form, which can bring essential difficulty in studying existence and asymptotic behavior of solutions. It seems impossible to reduce the second-order operator \mathcal{L}_H to the standard Laplace operator by means of a single change of coordinates. We

will give rigorous justification of the relation between the coefficient matrix K_H and limiting location, energy and the concentration diameter of the ground states, which is totally different from the 2D and 3D axisymmetric cases.

To state our results, we need to introduce some notations first. Let $\kappa(w) = \int_{\Omega} w(x) dx$ be the circulation of the vorticity w. For two sets A, B, we define $dist(A, B) = \min_{x \in A, y \in B} |x - B|$ y the distance between sets A and B and diam(A) the diameter of the set A.

We first consider desingularization of steady solutions of 3D Euler equations with helical symmetry with impermeable boundary condition $v_n \equiv 0$. Since (w, φ) is a steady solution, that is, the distribution of w, φ is independent of t, by (1.16), (w, φ) satisfies the steady vorticity equations

$$\begin{cases}
\nabla w \cdot \nabla^{\perp} \varphi = 0, \\
w = \mathcal{L}_H \varphi, \\
\varphi|_{\partial \Omega} = 0.
\end{cases}$$
(1.17)

Formally, if

$$\mathcal{L}_H \varphi = w = \frac{1}{\varepsilon^2} f(\varphi - \mu), \quad \varphi|_{\partial\Omega} = 0,$$

for some function f and constants ε, μ , then (1.17) automatically holds. To conclude, it suffices to look for solutions of the semilinear elliptic equations in divergence form

$$\begin{cases} \mathcal{L}_{H}\varphi(x) = -\operatorname{div} \cdot (K_{H}(x)\nabla\varphi(x)) = \frac{1}{\varepsilon^{2}}f(\varphi(x) - \mu), & x \in \Omega, \\ \varphi(x) = 0, & x \in \partial\Omega. \end{cases}$$
(1.18)

Our first result is as follows.

Theorem 1.2. For every $k > 0, m > 0, 0 < \varepsilon < 1$, there exists a family of helical solution pairs $(\mathbf{v}_{\varepsilon}, P_{\varepsilon})(x, t) \in C^1(D \times \mathbb{R}^+)$ of Euler equations (1.1) such that the support set of $\operatorname{curl} \mathbf{v}_{\varepsilon}$ is a topological helical tube and the associated vorticity-stream function pair $(w_{\varepsilon}, \varphi_{\varepsilon})$ is a solution of steady vorticity equations (1.17). Moreover, there holds

- (1) $\mathbf{v}_{\varepsilon} \cdot \mathbf{n} = 0 \text{ on } \partial D$.
- (2) The support set of w_{ε} is simply-connected and

$$\lim_{\varepsilon \to 0} \frac{\ln diam(supp(w_{\varepsilon}))}{\ln \varepsilon} = 1.$$

- As a consequence, $\lim_{\varepsilon \to 0} diam(supp(w_{\varepsilon})) = 0$. (3) $\lim_{\varepsilon \to 0} dist(supp(w_{\varepsilon}), x^*) = 0$, where $x^* \in \overline{\Omega}$ satisfies $|x^*| = \max_{\overline{\Omega}} |x|$.
- (4) $\lim_{\varepsilon \to 0} \kappa(w_{\varepsilon}) = \frac{2k\pi m}{\sqrt{k^2 + |x^*|^2}}.$

The solution is constructed by studying the existence and asymptotic behavior of the ground state solutions φ_{ε} of equations (1.18) with $f(t) = t_{+}^{p}$ for p > 1, $\mu = m \ln \frac{1}{\varepsilon}$ for some prescribed constant m > 0.

Remark 1.3. In [10], Dávila et al. constructed rotational-invariant solutions of vorticity equations with angular velocity $\alpha |\ln \varepsilon|$ in \mathbb{R}^2 . However, because of the choice of f_{ε} , the support set of vorticity is still the whole plane. Our result shows the existence of a family of steady solutions of (1.16) in a general bounded domain, such that the corresponding vorticity has non-vanishing circulation with small cross-section and shrinks to a helical filament as $\varepsilon \to 0$.

Remark 1.4. By the physical meaning of k, the sign of k determines two different helical structure. The cases k > 0 and k < 0 correspond to the left-handed helical structure and right-handed helical structure, respectively. For the case k < 0, one can similarly get solutions of (1.18) concentrating near a single point.

Our second result is on the desingularization of steady solutions of vorticity equations when the boundary is penetrable. Assume that $\mathbf{v} \cdot \mathbf{n} = v_n \ln \frac{1}{\varepsilon}$ for some helical function $v_n \not\equiv 0$. By (1.15), steady solution pairs (w, φ) satisfy

$$\begin{cases}
\nabla w \cdot \nabla^{\perp} \varphi = 0, \\
w = \mathcal{L}_{H} \varphi, \\
\nabla^{\perp} \varphi \cdot \nu|_{\partial \Omega} = v_{n}|_{\partial \Omega} \ln \frac{1}{\varepsilon}.
\end{cases}$$
(1.19)

Suppose that $q \in C^2(\Omega) \cap C^1(\overline{\Omega})$ satisfies

$$\begin{cases} \mathcal{L}_H q = 0, \\ \nabla^{\perp} q \cdot \nu|_{\partial\Omega} = -v_n|_{\partial\Omega}. \end{cases}$$
 (1.20)

Note that for a solution q of (1.20), q + C is also a solution for any constant C. Thus one can always assume that $\min_{\overline{\Omega}} q > 0$. Let $u = \varphi + q \ln \frac{1}{\varepsilon}$. Then u satisfies

$$\begin{cases}
\nabla w \cdot \nabla^{\perp} (u - q \ln \frac{1}{\varepsilon}) = 0, \\
w = \mathcal{L}_H u, \\
\nabla^{\perp} u \cdot \nu|_{\partial\Omega} = 0.
\end{cases}$$
(1.21)

So if

$$\mathcal{L}_H u = w = \frac{1}{\varepsilon^2} f(u - q \ln \frac{1}{\varepsilon}), \quad u|_{\partial\Omega} = 0,$$

for some function f and constants ε, μ , then (1.21) automatically holds. And the solution pairs (w, φ) of (1.19) can be obtained by letting $w = \mathcal{L}_H u$ and $\varphi = u - q \ln \frac{1}{\varepsilon}$.

Let $det(K_H)$ denote the determinant of K_H . Our second result is as follows.

Theorem 1.5. Let k > 0, q > 0 satisfy $\mathcal{L}_H q = 0$ and v_n be a helical function defined on ∂D with $v_n|_{\partial\Omega} = -\nabla^\perp q \cdot \nu|_{\partial\Omega}$. Then for every $0 < \varepsilon < 1$, there exists a family of helical solution pairs $(\mathbf{v}_{\varepsilon}, P_{\varepsilon})(x, t) \in C^1(D \times \mathbb{R}^+)$ of Euler equations (1.1) such that the support set of curl \mathbf{v}_{ε} is a topological helical tube and the associated vorticity-stream function pair $(w_{\varepsilon}, \varphi_{\varepsilon})$ is a solution of steady vorticity equations (1.19). Moreover, the following conclusions hold

- (1) $\mathbf{v}_{\varepsilon} \cdot \mathbf{n} = v_n \ln \frac{1}{\varepsilon}$ on ∂D .
- (2) The support set of w_{ε} is simply-connected and

$$\lim_{\varepsilon \to 0} \frac{\ln diam(supp(w_{\varepsilon}))}{\ln \varepsilon} = 1.$$

(3) $\lim_{\begin{subarray}{c} \varepsilon \to 0 \\ is. \end{subarray}} dist(supp(w_{\varepsilon}), x^*) = 0$, where $x^* \in \overline{\Omega}$ is a minimum point of $q^2 \sqrt{\det(K_H)}$, that is.

$$q(x^*)^2 \sqrt{\det(K_H(x^*))} = \min_{\overline{O}} q^2 \sqrt{\det(K_H)}.$$

- (4) $\lim_{\varepsilon \to 0} \int_{\Omega} w_{\varepsilon}(x) dx = 2\pi q(x^*) \sqrt{\det(K_H(x^*))}$.
- (5) Moreover, if $x^* \in \Omega$, then there exist $R_1, R_2 > 0$ satisfying

$$R_1 \varepsilon \leq diam(supp(w_{\varepsilon})) \leq R_2 \varepsilon.$$

We now give some comments of the proof of Theorem 1.5. The strategy is to consider the existence and limiting behavior of ground state solutions of a semilinear elliptic equation in divergence form, see (2.1) in section 2. First, using the critical point theory the existence of mountain pass solutions u_{ε} of (2.1) with critical value c_{ε} of the corresponding variational functional is proved. Then by choosing proper test functions, we get the upper bound of c_{ε} , from which we get the connectness of the vortex core. Since K_H is not $-\Delta$, one can not use the classical test functions to get accurate upper bounded of c_{ε} . The boundedness of the energy of the vortex core is then obtained. Finally based on the classical estimates of capacity (see, e.g., [3, 11, 24]), we prove the lower bound of c_{ε} and the limiting location of the core. The key of proof is to show that the concentration point of ground states u_{ε} is a minimum point of $q^2\sqrt{\det(K_H)}$. To this end, the optimal upper and lower bounds of c_{ε} must be obtained.

Remark 1.6. The results of Theorem 1.5 can be regarded as a general result of the desingularization of classical planar vortex case (see [20, 24]) and the vortex ring case (see [11]). Note that the cases of planar vortices and vortex rings correspond to the coefficient matrix $K_H(x) = Id$ and $\frac{1}{r}Id$, respectively. In [11], by considering solutions of

$$\begin{cases} -\operatorname{div}\left(\frac{1}{b}\nabla u\right) = \frac{1}{\varepsilon^2}b\left(u - q\ln\frac{1}{\varepsilon}\right)_+^{p-1}, & x \in \Omega, \\ u = 0, & x \in \partial\Omega, \end{cases}$$

where b is a scalar function and q is a positive function satisfying $-\operatorname{div}(\frac{1}{b}\nabla q)=0$, the authors constructed a family of C^1 solutions u_{ε} with nonvanishing circulation concentrating near a minimizer of q^2/b as $\varepsilon \to 0$. Indeed, if we choose $K_H(x)=\frac{1}{b}Id$ in Theorem 1.5, then solutions will concentrate near minimizers of $q^2\sqrt{\det(K_H)}=q^2/b$, which coincides with the results in [11].

Remark 1.7. Indeed, the existence of solutions of general elliptic equations in divergence form has been studied by [23], who considered a singularly perturbed elliptic problem:

$$-\varepsilon^2 \operatorname{div}(K(x)\nabla u) + V(x)u = u^p, \quad x \in \mathbb{R}^n, \tag{1.22}$$

where K(x) is strictly positive definite, $n \geq 3$, $p \in (1, \frac{n+2}{n-2})$ and $V \in C^1(\mathbb{R}^n)$ is positive. The authors constructed solutions concentrating near minimizers of $V(x)^{\frac{p+1}{p-1}}\sqrt{\det(K(x))}$ by the penalization technique. However, it seems that the method in [23] can not be used in our situation since it depends on the positiveness of V.

Remark 1.8. Recently, [8] considered desingularization of rotational-invariant solutions of 3D incompressible Euler equation with helical symmetry in an infinite pipe. Using properties of Green's function of a general uniformly elliptic operator, [8] proved the existence of mountain pass solutions of Euler equation with helical symmetry, the associated vorticities of which are rotational-invariant and concentrate near a helix. While in this paper, instead of using properties of Green's function, we use the estimates of capacity to improve estimates of the diameter of the vortex core and the energy of ground states in [8].

This paper is organized as follows. In section 2, we introduce the associated variational structure and prove the existence of mountain pass solutions of (2.1) for every $\varepsilon \in (0, 1)$. Some fundamental properties which will be used in section 3 are also proved. In section 3 we prove the limiting behavior of u_{ε} . The proof of Theorem 1.5 and Theorem 1.2 will be given in section 4.

2. Variational problem

We now consider the following equations

$$\begin{cases} \mathcal{L}_H u = -\operatorname{div}(K_H(x)\nabla u) = \frac{1}{\varepsilon^2} \left(u - q \ln \frac{1}{\varepsilon} \right)_+^p, & x \in \Omega, \\ u = 0, & x \in \partial\Omega, \end{cases}$$
 (2.1)

where p > 1, $\Omega \subseteq \mathbb{R}^2$ is bounded, $K_H(x) = \frac{1}{k^2 + x_1^2 + x_2^2} \begin{pmatrix} k^2 + x_2^2 & -x_1 x_2 \\ -x_1 x_2 & k^2 + x_1^2 \end{pmatrix}$ and q is a function defined in $\overline{\Omega}$ satisfying

(Q1). $q \in C^2(\Omega) \cap C^1(\overline{\Omega})$ and q > 0 in $\overline{\Omega}$.

(Q2). q is a \mathcal{L}_H - harmonic function, i.e., $\mathcal{L}_H q = -\text{div}(K_H(x)\nabla q) = 0$.

Let $(K_H(x)\mathbf{a}|\mathbf{b}) = \sum_{i,j=1}^2 (K_H)_{i,j}(x)a_ib_j$ for two vectors \mathbf{a}, \mathbf{b} . Define

$$\mathcal{H}(\Omega) = \left\{ u \in H_0^1(\Omega) \mid \int_{\Omega} (K_H(x) \nabla u | \nabla u) dx < +\infty \right\}$$

with the norm

$$||u||_{\mathcal{H}(\Omega)} := \left(\int_{\Omega} (K_H(x)\nabla u | \nabla u) dx \right)^{\frac{1}{2}}.$$

Since K_H is a positive definite matrix with two positive eigenvalues $\lambda_1 = 1$ and $\lambda_2 = \frac{k^2}{k^2 + x_1^2 + x_2^2}$, two norms $||\cdot||_{\mathcal{H}(\Omega)}$ and $||\cdot||_{H_0^1(\Omega)}$ are equivalent.

Define the associated energy functional of (2.1)

$$I_{\varepsilon}(u) = \frac{1}{2} \int_{\Omega} (K_H(x) \nabla u | \nabla u) dx - \frac{1}{(p+1)\varepsilon^2} \int_{\Omega} \left(u - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx, \quad \forall u \in \mathcal{H}(\Omega).$$
 (2.2)

By the definition of \mathcal{H} , I_{ε} is a well-defined C^1 functional on \mathcal{H} .

Define the Nehari manifold

$$\mathcal{N}_{\varepsilon} = \{ u \in \mathcal{H}(\Omega) \setminus \{0\} \mid \langle I'_{\varepsilon}(u), u \rangle = 0 \}$$

$$= \left\{ u \in \mathcal{H}(\Omega) \setminus \{0\} \mid \int_{\Omega} (K_H(x) \nabla u | \nabla u) dx = \frac{1}{\varepsilon^2} \int_{\Omega} \left(u - q \ln \frac{1}{\varepsilon} \right)_+^p u dx \right\}.$$
(2.3)

2.1. **Existence of solutions.** First, using the classical critical point theory, we get ground state solutions of (2.1).

Since the nonlinearity $f(t,x) = \frac{1}{\varepsilon^2} ((t - q(x) \ln \frac{1}{\varepsilon})_+^p)$ for p > 1, $I_{\varepsilon}(u)$ has a mountain pass geometry. Thus we can define the mountain pass value

$$c_{\varepsilon} = \inf_{\gamma \in \mathcal{P}_{\varepsilon}} \max_{t \in [0,1]} I_{\varepsilon}(\gamma(t)),$$

where

$$\mathcal{P}_{\varepsilon} = \{ \gamma \in C([0,1], \mathcal{H}(\Omega)) \mid \gamma(0) = 0, I_{\varepsilon}(\gamma(1)) < 0 \}.$$

Clearly, $c_{\varepsilon} > 0$. We have the following characterization of $\mathcal{N}_{\varepsilon}$ and the mountain pass value c_{ε} , see [7, 16].

Lemma 2.1 ([7], Theorem 1.3.7). For any $u \in \mathcal{N}_{\varepsilon}$, $u_{+} \not\equiv 0$. For any $u \in \mathcal{H}(\Omega)$ with $u_{+} \not\equiv 0$, there exists a unique t(u) > 0 such that $t(u)u \in \mathcal{N}_{\varepsilon}$. The value of t(u) is characterized by the identity

$$I_{\varepsilon}(t(u)u) = \max\{I_{\varepsilon}(tu), t > 0\}. \tag{2.4}$$

Moreover, there holds

$$c_{\varepsilon} \leq \inf_{w \neq 0, w \in \mathcal{H}(\Omega)} \max_{t \geq 0} I_{\varepsilon}(tw) = \inf_{w \in \mathcal{N}_{\varepsilon}} I_{\varepsilon}(w).$$

Finally, if the mountain pass value c_{ε} is a critical value for I_{ε} , then $c_{\varepsilon} = \inf_{w \in \mathcal{N}_{\varepsilon}} I_{\varepsilon}(w)$ is the least nontrivial critical value.

Using the mountain pass theorem, we get the existence of mountain pass solutions of I_{ε} with c_{ε} .

Proposition 2.2. I_{ε} has a mountain pass solution with mountain pass value c_{ε} . Namely, one can find $u_{\varepsilon} \in \mathcal{H}(\Omega)$ satisfying

$$I_{\varepsilon}(u_{\varepsilon}) = c_{\varepsilon}, \quad I'_{\varepsilon}(u_{\varepsilon}) = 0.$$
 (2.5)

As a consequence, there holds $c_{\varepsilon} = \inf_{w \not\equiv 0, w \in \mathcal{H}(\Omega)} \max_{t \geq 0} I_{\varepsilon}(tw) = \inf_{w \in \mathcal{N}_{\varepsilon}} I_{\varepsilon}(w)$.

Proof. By standard mountain pass theory (see, e.g., §1.4 in [25]), we can prove that there exists u_{ε} such that $I'_{\varepsilon}(u_{\varepsilon}) = 0$ and $I_{\varepsilon}(u_{\varepsilon}) = c_{\varepsilon}$. So by Lemma 2.1, $c_{\varepsilon} = \inf_{w \in \mathcal{N}_{\varepsilon}} I_{\varepsilon}(w)$.

2.2. **Basic properties.** First, we give some basic properties of I_{ε} and the operator \mathcal{L}_H as follows.

Lemma 2.3. For any $u \in \mathcal{H}(\Omega)$,

$$\left(\frac{1}{2} - \frac{1}{p+1}\right) ||u||_{\mathcal{H}(\Omega)}^2 \le I_{\varepsilon}(u) - \frac{1}{p+1} \langle I_{\varepsilon}'(u), u \rangle. \tag{2.6}$$

Proof. It follows from the definition of I_{ε} and I'_{ε} that

$$I_{\varepsilon}(u) - \frac{1}{p+1} \langle I'_{\varepsilon}(u), u \rangle$$

$$= \frac{1}{2} ||u||^{2}_{\mathcal{H}(\Omega)} - \frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(u - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx - \left(\frac{1}{p+1} ||u||^{2}_{\mathcal{H}(\Omega)} - \frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(u - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} u dx \right)$$

$$\geq \left(\frac{1}{2} - \frac{1}{p+1} \right) ||u||^{2}_{\mathcal{H}(\Omega)}.$$

Lemma 2.4. For any $u \in \mathcal{H}(\Omega)$,

$$\int_{\Omega} (K_H(x)\nabla u|\nabla u)dx = \int_{\Omega} q^2 \left(K_H(x)\nabla\left(\frac{u}{q}\right)|\nabla\left(\frac{u}{q}\right)\right)dx. \tag{2.7}$$

Proof. We first claim that

$$\left(K_H(x)\nabla q|\nabla\left(\frac{u^2}{q}\right)\right) = \left(K_H(x)\nabla u|\nabla u\right) - q^2\left(K_H(x)\nabla\left(\frac{u}{q}\right)|\nabla\left(\frac{u}{q}\right)\right).$$
(2.8)

Indeed, we have

$$(K_{H})_{11}(\partial_{1}u)^{2} - q^{2}(K_{H})_{11} \left(\partial_{1}\left(\frac{u}{q}\right)\right)^{2} = (K_{H})_{11} \left(\frac{2u}{q}\partial_{1}u\partial_{1}q - \frac{u^{2}}{q^{2}}(\partial_{1}q)^{2}\right) = (K_{H})_{11}\partial_{1}q\partial_{1}\left(\frac{u^{2}}{q}\right),$$

$$((K_{H})_{12} + (K_{H})_{21})\partial_{1}u\partial_{2}u - q^{2}((K_{H})_{12} + (K_{H})_{21})\partial_{1}\left(\frac{u}{q}\right)\partial_{2}\left(\frac{u}{q}\right)$$

$$= 2(K_{H})_{12}\frac{u}{q}(\partial_{1}q\partial_{2}u + \partial_{2}q\partial_{1}u) - 2(K_{H})_{12}\frac{u^{2}}{q^{2}}\partial_{1}q\partial_{2}q$$

$$= (K_{H})_{12}\partial_{1}q\partial_{2}\left(\frac{u^{2}}{q}\right) + (K_{H})_{21}\partial_{2}q\partial_{1}\left(\frac{u^{2}}{q}\right),$$

and

$$(K_H)_{22}(\partial_2 u)^2 - q^2(K_H)_{22} \left(\partial_2 \left(\frac{u}{q} \right) \right)^2 = (K_H)_{22} \left(\frac{2u}{q} \partial_2 u \partial_2 q - \frac{u^2}{q^2} (\partial_2 q)^2 \right) = (K_H)_{22} \partial_2 q \partial_2 \left(\frac{u^2}{q} \right).$$

Adding up the above inequalities, we get (2.8).

Since $\mathcal{L}_H q = 0$ and $\frac{u^2}{q} \in H_0^1(\Omega)$, we have $\int_{\Omega} \left(K_H(x) \nabla q | \nabla \left(\frac{u^2}{q} \right) \right) dx = 0$. Integrating both sides of (2.8) over Ω , we get (2.7).

For any $\bar{x} = (|\bar{x}| \cos \theta_{\bar{x}}, |\bar{x}| \sin \theta_{\bar{x}}) \in \mathbb{R}^2$, denote $\bar{R}_{\bar{x}} = \begin{pmatrix} \cos \theta_{\bar{x}} & \sin \theta_{\bar{x}} \\ -\sin \theta_{\bar{x}} & \cos \theta_{\bar{x}} \end{pmatrix}$ the rotational transformation matrix through $\theta_{\bar{x}}$. Then we have

Lemma 2.5. For any $\bar{x} \in \mathbb{R}^2$, there holds

$$K_H(\bar{R}_{\bar{x}}y) = \bar{R}_{\bar{x}}K_H(y)\bar{R}_{\bar{x}}^t, \quad \forall y \in \mathbb{R}^2.$$

Proof. Note that for any $y = (y_1, y_2)^t$, $\bar{R}_{\bar{x}}y = (y_1 \cos \theta_{\bar{x}} + y_2 \sin \theta_{\bar{x}}, y_2 \cos \theta_{\bar{x}} - y_1 \sin \theta_{\bar{x}})^t$ and $|\bar{R}_{\bar{x}}y| = |y|$. By the definition of K_H , we get

$$K_{H}(\bar{R}_{\bar{x}}y) = \frac{1}{k^{2} + |y|^{2}} \begin{pmatrix} k^{2} + (y_{2}\cos\theta_{\bar{x}} - y_{1}\sin\theta_{\bar{x}})^{2} & -(y_{1}\cos\theta_{\bar{x}} + y_{2}\sin\theta_{\bar{x}})(y_{2}\cos\theta_{\bar{x}} - y_{1}\sin\theta_{\bar{x}}) \\ -(y_{1}\cos\theta_{\bar{x}} + y_{2}\sin\theta_{\bar{x}})(y_{2}\cos\theta_{\bar{x}} - y_{1}\sin\theta_{\bar{x}}) & k^{2} + (y_{1}\cos\theta_{\bar{x}} + y_{2}\sin\theta_{\bar{x}})^{2} \end{pmatrix}$$

$$= \frac{1}{k^{2} + |y|^{2}} \begin{pmatrix} \cos\theta_{\bar{x}} & \sin\theta_{\bar{x}} \\ -\sin\theta_{\bar{x}} & \cos\theta_{\bar{x}} \end{pmatrix} \begin{pmatrix} k^{2} + y_{2}^{2} & -y_{1}y_{2} \\ -y_{1}y_{2} & k^{2} + y_{1}^{2} \end{pmatrix} \begin{pmatrix} \cos\theta_{\bar{x}} & -\sin\theta_{\bar{x}} \\ \sin\theta_{\bar{x}} & \cos\theta_{\bar{x}} \end{pmatrix}$$

$$= \bar{R}_{\bar{x}}K_{H}(y)\bar{R}_{\bar{x}}^{t}.$$

A direct consequence of Lemma 2.5 is the rotational invariance of the problem (2.1), which will be used in the proof of Theorem 1.5 in section 4. Define $\Omega_{\bar{x}} := \{\bar{R}_{\bar{x}}x \mid x \in \Omega\}$ the region of Ω rotated clockwise by $\theta_{\bar{x}}$. For a function $u \in \mathcal{H}(\Omega)$, we define

$$u_{\bar{x}}(x) := u(\bar{R}_{-\bar{x}}x) \quad \forall x \in \Omega_{\bar{x}}.$$

So $u_{\bar{x}} \in \mathcal{H}(\Omega_{\bar{x}})$. Define $q_{\bar{x}}(x) := q(\bar{R}_{-\bar{x}}x)$ for any $x \in \Omega_{\bar{x}}$. Then we get

Lemma 2.6. [Rotational invariance] u is a solution of (2.1) if and only if $u_{\bar{x}} \in \mathcal{H}(\Omega_{\bar{x}})$ satisfies

$$-\operatorname{div}(K_H(x)\nabla u_{\bar{x}}) = \frac{1}{\varepsilon^2} \left(u_{\bar{x}} - q_{\bar{x}} \ln \frac{1}{\varepsilon} \right)_+^p, \quad \text{in } \Omega_{\bar{x}}.$$
 (2.9)

Proof. For any $x \in \Omega_{\bar{x}}$, let $y = \bar{R}_{-\bar{x}}x$. It is not hard to check that for any function g and vector field $\mathbf{F} = (f_1, f_2)^t$, there holds

$$\nabla_x g(\bar{R}_{-\bar{x}}x) = \bar{R}_{\bar{x}}\nabla_y g(y),$$

and

$$\nabla_x \cdot \mathbf{F}(\bar{R}_{-\bar{x}}x) = \nabla_y \cdot (\bar{R}_{-\bar{x}}\mathbf{F})(y).$$

So by using Lemma 2.5, we get

$$-\operatorname{div}(K_{H}(x)\nabla u_{\bar{x}}(x)) = -\nabla_{x} \cdot (K_{H}(x)\nabla_{x}u(\bar{R}_{-\bar{x}}x))$$

$$= -\nabla_{x} \cdot (K_{H}(x)(\bar{R}_{\bar{x}}\nabla_{y}u)(\bar{R}_{-\bar{x}}x))$$

$$= -\nabla_{x} \cdot (\bar{R}_{\bar{x}}K_{H}(\bar{R}_{-\bar{x}}x)\bar{R}_{\bar{x}}^{t}(\bar{R}_{\bar{x}}\nabla_{y}u)(\bar{R}_{-\bar{x}}x))$$

$$= -\nabla_{y} \cdot (\bar{R}_{-\bar{x}}\bar{R}_{\bar{x}}K_{H}(y)\nabla_{y}u(y))$$

$$= -\nabla_{y} \cdot (K_{H}(y)\nabla_{y}u)(y)$$

$$= \frac{1}{\varepsilon^{2}} \left(u(y) - q(y)\ln\frac{1}{\varepsilon}\right)_{+}^{p}$$

$$= \frac{1}{\varepsilon^{2}} \left(u_{\bar{x}}(x) - q_{\bar{x}}(x)\ln\frac{1}{\varepsilon}\right)_{+}^{p}.$$

3. Asymptotic behavior of u_{ε}

Now, we give the asymptotic behavior of mountain pass solutions u_{ε} of (2.1). We first consider the asymptotic behavior of u_{ε} under an extra assumption of q and K_H :

 (A_1) . There exist minimum points of $q^2\sqrt{det(K_H)}$ over $\overline{\Omega}$ which is on the x_1 -axis.

Indeed, this additional assumption is not essential. Under this assumption, it is convenient to give an optimal upper bound of c_{ε} , see Proposition 3.1. The proof of asymptotic behavior of u_{ε} without the assumption (A_1) will be given in section 4.

Let
$$x^* \in \overline{\Omega} \cap \{x = (x_1, x_2) \mid x_2 = 0\}$$
 be such that

$$q^{2}(x^{*})\sqrt{\det(K_{H}(x^{*}))} = \min_{x \in \overline{\Omega}} q^{2}(x)\sqrt{\det(K_{H}(x))}.$$

Let $0 < \varepsilon < 1$.

3.1. **Upper bound of** c_{ε} . First, we compute the upper bound of c_{ε} . By choosing proper competitors, we can get the following upper bound of c_{ε} .

Proposition 3.1. There holds

$$\limsup_{\varepsilon \to 0} \frac{c_{\varepsilon}}{\ln \frac{1}{\varepsilon}} \le \pi q^2 \sqrt{\det(K_H)}(x^*) = \pi \min_{x \in \overline{\Omega}} q^2 \sqrt{\det(K_H)}(x).$$

Moreover, if $x^* \in \Omega$, then

$$c_{\varepsilon} \le \pi \min_{x \in \overline{\Omega}} q^2 \sqrt{\det(K_H)}(x) \ln \frac{1}{\varepsilon} + O(1).$$

Proof. Let U(x) be a C^{∞} radially symmetric function such that

$$\begin{cases} U(x) \ge 0, & x \in B_1(0), \\ U(x) = \ln \frac{1}{|x|}, & x \in B_1(0)^c. \end{cases}$$

For any $\bar{x} \in \Omega \cap \{x_2 = 0\}$, we choose $\delta > 0$ sufficiently small and a truncation $\varphi_{\delta} \in C_c^{\infty}(B_{2\delta}(0))$ such that

$$0 \le \phi_{\delta} \le 1$$
 in $B_{2\delta}(0)$; $\phi_{\delta} = 1$ in $B_{\delta}(0)$.

For any constants $l_1, l_2 > 0$ (which will be determined later), define

$$\hat{U}(x_1, x_2) = U\left(\frac{x_1}{l_1}, \frac{x_2}{l_2}\right).$$

So the support set of \hat{U}_+ is the ellipse $\left\{(x_1,x_2)\mid \frac{x_1^2}{l_1^2}+\frac{x_2^2}{l_2^2}\leq 1\right\}$. For any set A, define $\hat{A}=\left\{(x_1,x_2)\mid \left(\frac{x_1}{l_1},\frac{x_2}{l_2}\right)\in A\right\}$. Let $\hat{\varphi}_{\delta}(x_1,x_2)=\varphi_{\delta}\left(\frac{x_1}{l_1},\frac{x_2}{l_2}\right)$. Then $supp(\hat{\varphi}_{\delta})=\hat{B}_{2\delta}$ and $\hat{\varphi}_{\delta}\equiv 1$ on \hat{B}_{δ} .

We define for any $\tau > 0$ a test function

$$v_{\varepsilon}^{\tau} = q(x) \left(\hat{U} \left(\frac{x - \bar{x}}{\varepsilon} \right) + \ln \frac{\tau}{\varepsilon} \right) \hat{\varphi}_{\delta}(x - \bar{x}) = \left[q \ln \frac{1}{\varepsilon} + q \left(\hat{U} \left(\frac{x - \bar{x}}{\varepsilon} \right) + \ln \tau \right) \right] \hat{\varphi}_{\delta}(x - \bar{x}).$$

Then $v_{\varepsilon}^{\tau} \in H_0^1(\Omega)$. Define

$$g_{\varepsilon}(\tau) := \frac{1}{\ln \frac{1}{\varepsilon}} \langle I_{\varepsilon}'(v_{\varepsilon}^{\tau}), v_{\varepsilon}^{\tau} \rangle = \frac{1}{\ln \frac{1}{\varepsilon}} \left(\int_{\Omega} (K_H(x) \nabla v_{\varepsilon}^{\tau} | \nabla v_{\varepsilon}^{\tau}) dx - \frac{1}{\varepsilon^2} \int_{\Omega} \left(v_{\varepsilon}^{\tau} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} v_{\varepsilon}^{\tau} dx \right).$$

We now prove that there exists $\tau_{\varepsilon} > 0$ such that $g_{\varepsilon}(\tau_{\varepsilon}) = 0$, that is, $v_{\varepsilon}^{\tau_{\varepsilon}} \in \mathcal{N}_{\varepsilon}$. By Lemma 2.4,

$$\int_{\Omega} (K_{H}(x)\nabla v_{\varepsilon}^{\tau}|\nabla v_{\varepsilon}^{\tau})dx = \int_{\Omega} q^{2} \left(K_{H}(x)\nabla \left(\frac{v_{\varepsilon}^{\tau}}{q}\right)|\nabla \left(\frac{v_{\varepsilon}^{\tau}}{q}\right)\right)dx
= \int_{\hat{B}_{2\delta}(\bar{x})} q^{2} \left(K_{H}(x)\nabla \left(\frac{v_{\varepsilon}^{\tau}}{q}\right)|\nabla \left(\frac{v_{\varepsilon}^{\tau}}{q}\right)\right)dx
= \left(\int_{\hat{B}_{2\delta}(\bar{x})/\hat{B}_{\delta}(\bar{x})} + \int_{\hat{B}_{\delta}(\bar{x})/\hat{B}_{\varepsilon}(\bar{x})} + \int_{\hat{B}_{\varepsilon}(\bar{x})}\right)q^{2} \left(K_{H}(x)\nabla \left(\frac{v_{\varepsilon}^{\tau}}{q}\right)|\nabla \left(\frac{v_{\varepsilon}^{\tau}}{q}\right)\right)dx
= A_{1} + A_{2} + A_{3}.$$

By the definition of v_{ε}^{τ} , we have

$$A_1 \le C \left(1 + \left| \ln \frac{\tau}{\delta} \right| \right), \tag{3.1}$$

where C is independent of τ . Since ε sufficiently small, we can assume that $\varepsilon < \delta$, which implies that

$$A_{3} = \int_{\hat{B}_{\varepsilon}(\bar{x})} q^{2} \left(K_{H}(x) \nabla \left(\frac{v_{\varepsilon}^{\tau}}{q} \right) | \nabla \left(\frac{v_{\varepsilon}^{\tau}}{q} \right) \right) dx = \int_{\hat{B}_{\varepsilon}(\bar{x})} q^{2} \left(K_{H}(x) \nabla \hat{U} \left(\frac{x - \bar{x}}{\varepsilon} \right) | \nabla \hat{U} \left(\frac{x - \bar{x}}{\varepsilon} \right) \right) dx$$

$$= \int_{\hat{B}_{1}(0)} q^{2} (\varepsilon y + \bar{x}) (K_{H}(\varepsilon y + \bar{x}) \nabla \hat{U}(y) | \nabla \hat{U}(y)) dy$$

$$\Rightarrow \int_{\hat{B}_{1}(0)} q^{2} (\bar{x}) (K_{H}(\bar{x}) \nabla \hat{U}(y) | \nabla \hat{U}(y)) dy \quad \text{as } \varepsilon \to 0.$$

$$(3.2)$$

The convergence is uniformly about τ .

As for A_2 , we have

$$A_{2} = \int_{\hat{B}_{\delta}(\bar{x})\backslash\hat{B}_{\varepsilon}(\bar{x})} q^{2} \left(K_{H}(x) \nabla \left(\frac{v_{\varepsilon}^{\tau}}{q} \right) | \nabla \left(\frac{v_{\varepsilon}^{\tau}}{q} \right) \right) dx$$

$$= \int_{\hat{B}_{\delta}(0)\backslash\hat{B}_{\varepsilon}(0)} q^{2} (x + \bar{x}) \left(K_{H}(x + \bar{x}) \nabla \hat{U} \left(\frac{x}{\varepsilon} \right) | \nabla \hat{U} \left(\frac{x}{\varepsilon} \right) \right) dx.$$

Note that for any vector $a = (a_1, a_2)$,

$$(K_H(x)a|a) = (K_H)_{11}(x)a_1^2 + ((K_H)_{12} + (K_H)_{21})(x)a_1a_2 + (K_H)_{22}(x)a_2^2.$$

Hence we have

$$A_{2} = \int_{\hat{B}_{\delta}(0)\backslash\hat{B}_{\varepsilon}(0)} q^{2}(K_{H})_{11}(x+\bar{x})\partial_{1}\hat{U}\left(\frac{x}{\varepsilon}\right)\partial_{1}\hat{U}\left(\frac{x}{\varepsilon}\right)dx$$

$$+ \int_{\hat{B}_{\delta}(0)\backslash\hat{B}_{\varepsilon}(0)} q^{2}((K_{H})_{12} + (K_{H})_{21})(x+\bar{x})\partial_{1}\hat{U}\left(\frac{x}{\varepsilon}\right)\partial_{2}\hat{U}\left(\frac{x}{\varepsilon}\right)dx$$

$$+ \int_{\hat{B}_{\delta}(0)\backslash\hat{B}_{\varepsilon}(0)} q^{2}(K_{H})_{22}(x+\bar{x})\partial_{2}\hat{U}\left(\frac{x}{\varepsilon}\right)\partial_{2}\hat{U}\left(\frac{x}{\varepsilon}\right)dx.$$

Note that $\nabla \hat{U}(x_1, x_2) = \left(-\frac{1}{l_1^2} \frac{x_1}{\frac{x_1^2}{l_1^2} + \frac{x_2^2}{l_2^2}}, -\frac{1}{l_2^2} \frac{x_2}{\frac{x_1^2}{l_1^2} + \frac{x_2^2}{l_2^2}}\right)$ on $\hat{B}_1(0)^c$. Hence direct calculation yields

$$\int_{\hat{B}_{\delta}(0)\backslash\hat{B}_{\varepsilon}(0)} q^{2}(K_{H})_{11}(x+\bar{x})\partial_{1}\hat{U}\left(\frac{x}{\varepsilon}\right)\partial_{1}\hat{U}\left(\frac{x}{\varepsilon}\right)dx
= \int_{\hat{B}_{\frac{\delta}{\varepsilon}}(0)\backslash\hat{B}_{1}(0)} q^{2}(K_{H})_{11}(\varepsilon x+\bar{x})\partial_{1}\hat{U}(x)\partial_{1}\hat{U}(x)dx
= \frac{1}{l_{1}^{4}} \int_{\hat{B}_{\frac{\delta}{\varepsilon}}(0)\backslash\hat{B}_{1}(0)} q^{2}(K_{H})_{11}(\varepsilon x+\bar{x}) \frac{x_{1}^{2}}{\left(\frac{x_{1}^{2}}{l_{1}^{2}}+\frac{x_{2}^{2}}{l_{2}^{2}}\right)^{2}} dx
= \frac{l_{2}}{l_{1}} \int_{B_{\frac{\delta}{\varepsilon}}(0)\backslash\hat{B}_{1}(0)} q^{2}(K_{H})_{11}(\varepsilon(l_{1}x_{1},l_{2}x_{2})+\bar{x}) \frac{x_{1}^{2}}{(x_{1}^{2}+x_{2}^{2})^{2}} dx
= \frac{l_{2}}{l_{1}} \int_{B_{\frac{\delta}{\varepsilon}}(0)\backslash\hat{B}_{1}(0)} (q^{2}(K_{H})_{11}(\bar{x})+\varepsilon\nabla(q^{2}(K_{H})_{11})(\bar{x})\cdot(l_{1}x_{1},l_{2}x_{2})+O(\varepsilon^{2}|x|^{2})) \frac{x_{1}^{2}}{(x_{1}^{2}+x_{2}^{2})^{2}} dx
= \frac{l_{2}}{l_{1}} \pi q^{2}(K_{H})_{11}(\bar{x}) \ln\frac{\delta}{\varepsilon}+O(1),$$
(3.3)

where we used $K_H, q \in \mathbb{C}^2$ and Taylor expansion. Similarly, we can get

$$\int_{\hat{B}_{\delta}(0)\backslash\hat{B}_{\varepsilon}(0)} q^{2}(K_{H})_{22}(x+\bar{x})\partial_{2}\hat{U}\left(\frac{x}{\varepsilon}\right)\partial_{2}\hat{U}\left(\frac{x}{\varepsilon}\right)dx = \frac{l_{1}}{l_{2}}\pi q^{2}(K_{H})_{22}(\bar{x})\ln\frac{\delta}{\varepsilon} + O(1). \quad (3.4)$$

Since \bar{x} is on the x_1 axis, we get $(K_H)_{12}(\bar{x}) = (K_H)_{21}(\bar{x}) = 0$, which implies that

$$\int_{\hat{B}_{\delta}(0)\backslash\hat{B}_{\varepsilon}(0)} q^{2}((K_{H})_{12} + (K_{H})_{21})(x + \bar{x})\partial_{1}\hat{U}\left(\frac{x}{\varepsilon}\right)\partial_{2}\hat{U}\left(\frac{x}{\varepsilon}\right)dx$$

$$= \int_{B_{\frac{\delta}{\varepsilon}}(0)\backslash B_{1}(0)} q^{2}((K_{H})_{12} + (K_{H})_{21})(\bar{x})\frac{x_{1}^{2}}{(x_{1}^{2} + x_{2}^{2})^{2}}dx + O(1)$$

$$= O(1). \tag{3.5}$$

So by (3.3), (3.4) and (3.5) we have

$$A_{2} = \frac{l_{2}}{l_{1}} \pi q^{2} (K_{H})_{11}(\bar{x}) \ln \frac{1}{\varepsilon} + \frac{l_{1}}{l_{2}} \pi q^{2} (K_{H})_{22}(\bar{x}) \ln \frac{1}{\varepsilon} + O(1), \tag{3.6}$$

where O(1) is some bounded quantity independent of τ .

Choosing $\frac{l_2}{l_1} = \sqrt{\frac{(K_H)_{22}}{(K_H)_{11}}}(\bar{x})$ and by (3.1), (3.2) and (3.6), we get

$$\int_{\Omega} (K_H(x)\nabla v_{\varepsilon}^{\tau}|\nabla v_{\varepsilon}^{\tau})dx = 2\pi q^2 \sqrt{(K_H)_{11}(K_H)_{22}}(\bar{x})\ln\frac{1}{\varepsilon} + O(1) + C|\ln\tau|
= 2\pi q^2 \sqrt{\det(K_H)}(\bar{x})\ln\frac{1}{\varepsilon} + O(1) + C|\ln\tau|,$$
(3.7)

from which we deduce that

$$\lim_{\varepsilon \to 0} \frac{1}{\ln \frac{1}{\varepsilon}} \int_{\Omega} (K_H(x) \nabla v_{\varepsilon}^{\tau} | \nabla v_{\varepsilon}^{\tau}) dx = 2\pi q^2 \sqrt{\det(K_H)}(\bar{x})$$
(3.8)

uniformly in any compact set of $\tau > 0$.

On the other hand, note that

$$\frac{1}{\varepsilon^2} \int_{\Omega} \left(v_{\varepsilon}^{\tau} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} v_{\varepsilon}^{\tau} dx = \frac{1}{\varepsilon^2} \int_{\Omega} \left(v_{\varepsilon}^{\tau} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx + \frac{1}{\varepsilon^2} \int_{\Omega} \left(v_{\varepsilon}^{\tau} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} q \ln \frac{1}{\varepsilon} dx.$$

By the definition of v_{ε}^{τ} , for ε sufficiently small and every $x \in \Omega$, we have

$$\left(v_{\varepsilon}^{\tau}(x) - q(x)\ln\frac{1}{\varepsilon}\right)_{+} = q(x)\left(\hat{U}\left(\frac{x - \bar{x}}{\varepsilon}\right) + \ln\tau\right)_{+}.$$

Hence we get

$$\frac{1}{\varepsilon^2} \int_{\Omega} \left(v_{\varepsilon}^{\tau} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx = \int_{\hat{B}_{\tau}(0)} q^{p+1} (\bar{x} + \varepsilon y) (\hat{U}(y) + \ln \tau)_{+}^{p+1} dy$$

$$\Rightarrow q^{p+1}(\bar{x}) \int_{\hat{B}_{\tau}(0)} (\hat{U}(y) + \ln \tau)_{+}^{p+1} dy,$$
(3.9)

and

$$\frac{1}{\varepsilon^2 \ln \frac{1}{\varepsilon}} \int_{\Omega} \left(v_{\varepsilon}^{\tau} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} q \ln \frac{1}{\varepsilon} dx \Rightarrow q^{p+1}(\bar{x}) \int_{\hat{B}_{\tau}(0)} (\hat{U}(y) + \ln \tau)_{+}^{p} dy. \tag{3.10}$$

The convergences are uniformly in any compact set of τ . By (3.9) and (3.10), we get

$$\lim_{\varepsilon \to 0} \frac{1}{\varepsilon^2 \ln \frac{1}{\varepsilon}} \int_{\Omega} \left(v_{\varepsilon}^{\tau} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} v_{\varepsilon}^{\tau} dx = q^{p+1}(\bar{x}) \int_{\hat{B}_{\tau}(0)} (\hat{U}(y) + \ln \tau)_{+}^{p} dy.$$
 (3.11)

It follows from (3.8) and (3.11) that for any $\tau > 0$, $\lim_{\varepsilon \to 0} g_{\varepsilon}(\tau) = g(\tau)$, where g is defined by $g(\tau) = 2\pi q^2 \sqrt{\det(K_H)}(\bar{x}) - q^{p+1}(\bar{x}) \int_{\hat{B}_{\tau}(0)} (\hat{U}(y) + \ln \tau)_+^p dy$, and the convergence is uniformly in any compact set of τ . Now it is not hard to prove that there exist two numbers $\tau_1, \tau_2 > 0$ such that $g(\tau_1) < 0 < g(\tau_2)$. So for ε sufficiently small, we have $g_{\varepsilon}(\tau_1) < 0 < g_{\varepsilon}(\tau_2)$, from which we deduce that, there exists $\tau_{\varepsilon} \in (\tau_1, \tau_2)$ satisfying $g_{\varepsilon}(\tau_{\varepsilon}) = 0$. Then $v_{\varepsilon}^{\tau_{\varepsilon}} \in \mathcal{N}_{\varepsilon}$.

Now by (3.8), (3.11) and $\tau_{\varepsilon} \in (\tau_1, \tau_2)$, we can compute that

$$\lim_{\varepsilon \to 0} \frac{1}{\ln \frac{1}{\varepsilon}} I_{\varepsilon}(v_{\varepsilon}^{\tau_{\varepsilon}}) = \lim_{\varepsilon \to 0} \frac{1}{2 \ln \frac{1}{\varepsilon}} \int_{\Omega} (K_{H}(x) \nabla v_{\varepsilon}^{\tau_{\varepsilon}} | \nabla v_{\varepsilon}^{\tau_{\varepsilon}}) dx - \lim_{\varepsilon \to 0} \frac{1}{(p+1)\varepsilon^{2} \ln \frac{1}{\varepsilon}} \int_{\Omega} \left(v_{\varepsilon}^{\tau_{\varepsilon}} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx$$
$$= \pi q^{2} \sqrt{\det(K_{H})}(\bar{x}).$$

Taking the infimum over $\bar{x} \in \Omega \cap \{x_2 = 0\}$ and using the assumption that x^* is on the x_1 axis, we get $\limsup_{\varepsilon \to 0} \frac{c_\varepsilon}{\ln \frac{1}{\varepsilon}} \le \pi q^2(x^*) \sqrt{\det(K_H(x^*))} = \pi \min_{x \in \overline{\Omega}} q^2 \sqrt{\det(K_H)}(x)$. If $x^* \in \Omega$, we can improve the above estimate as follows. Indeed, choosing $\bar{x} = x^*$, and

using (3.7) and (3.9) again, we can get

$$c_{\varepsilon} \leq I_{\varepsilon}(v_{\varepsilon}^{\tau_{\varepsilon}}) = \frac{1}{2} \int_{\Omega} (K_{H}(x) \nabla v_{\varepsilon}^{\tau_{\varepsilon}} | \nabla v_{\varepsilon}^{\tau_{\varepsilon}}) dx - \frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(v_{\varepsilon}^{\tau_{\varepsilon}} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx$$
$$= \pi q^{2} \sqrt{\det(K_{H})} (x^{*}) \ln \frac{1}{\varepsilon} + O(1),$$

where we used the fact that $\tau_{\varepsilon} \in (\tau_1, \tau_2)$ and the limit in (3.9) is uniform on compact sets of τ . The proof is thus complete.

3.2. Diameter and connectness of the vortex core. We now prove the connectness and estimate the diameter of the vortex core. To this end, we define A_{ε} the vortex core of solution u_{ε} , that is,

$$A_{\varepsilon} = \left\{ x \in \Omega \mid u_{\varepsilon}(x) > q(x) \ln \frac{1}{\varepsilon} \right\}.$$

Clearly by the classical regularity theory of elliptic equations, $u_{\varepsilon} \in C^{2,\alpha}(\Omega)$ for any $\alpha \in$ (0,1) and A_{ε} is an open subset of Ω .

 $\max_{x,y \in supp(A_{\varepsilon})} |x-y|$. We can prove the connectness and estimate the Define diam $(A_{\varepsilon}) =$ diameter of A_{ε} as follows.

Proposition 3.2. For every $\varepsilon > 0$ sufficiently small, A_{ε} is connected and simply connected. Moreover

$$\lim_{\varepsilon \to 0} \frac{diam(A_{\varepsilon})}{dist(A_{\varepsilon}, \partial \Omega)} = 0.$$

As a consequence, $diam(A_{\varepsilon})$ tends to 0 as $\varepsilon \to 0$.

Proof. Assume that A_{ε} has two components A_1, A_2 . We denote $\psi_i = \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon}\right) \chi_{A_i} \in$ $H_0^1(\Omega)$. Let $\eta_0 > 0$ to be determined later. Define $\tilde{w}_{\varepsilon}(s) = u_{\varepsilon} + s\psi_1 - s\eta_0\psi_2$. Then $\tilde{w}_{\varepsilon}(s) \in H_0^1(\Omega)$ for $s \geq 0$ sufficiently small. By Proposition 2.2, t_0 is a maximum point of $I_{\varepsilon}(t\tilde{w}_{\varepsilon})$ if and only if

$$t_0 \int_{\Omega} (K_H(x) \nabla \tilde{w}_{\varepsilon} | \nabla \tilde{w}_{\varepsilon}) dx = \frac{1}{\varepsilon^2} \int_{\Omega} \left(t_0 \tilde{w}_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_+^p \tilde{w}_{\varepsilon} dx.$$

Let $D(s,t) = t \int_{\Omega} (K_H(x) \nabla \tilde{w}_{\varepsilon} | \nabla \tilde{w}_{\varepsilon}) dx - \frac{1}{\varepsilon^2} \int_{\Omega} (t \tilde{w}_{\varepsilon} - q \ln \frac{1}{\varepsilon})_+^p \tilde{w}_{\varepsilon} dx$. Then D(0,1) = 0. By q > 0, we have

$$\frac{\partial D(0,t)}{\partial t}\Big|_{t=1} = \int_{\Omega} (K_H(x)\nabla u_{\varepsilon}|\nabla u_{\varepsilon})dx - \frac{p}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_{+}^{p-1} u_{\varepsilon}^2 dx
= \frac{1}{\varepsilon^2} \int_{\Omega} \left(\left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_{+}^{p} u_{\varepsilon} - p\left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_{+}^{p-1} u_{\varepsilon}^2\right) dx < 0.$$

Hence by the implicit function theorem, there is a function t = t(s) in the neighborhood of s = 0 such that t(0) = 1 and D(s, t(s)) = 0, which implies that, $t(s)\tilde{w}_{\varepsilon} \in \mathcal{N}_{\varepsilon}$. Note that

$$D_{s}(0,1) = 2 \int_{\Omega} (K_{H}(x) \nabla u_{\varepsilon} | \nabla (\psi_{1} - \eta_{0} \psi_{2})) dx - \frac{1}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} (\psi_{1} - \eta_{0} \psi_{2}) dx$$
$$- \frac{1}{\varepsilon^{2}} \int_{\Omega} p \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p-1} u_{\varepsilon} (\psi_{1} - \eta_{0} \psi_{2}) dx$$
$$= - \frac{1}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p-1} \left((p-1)u_{\varepsilon} + q \ln \frac{1}{\varepsilon} \right) (\psi_{1} - \eta_{0} \psi_{2}) dx.$$

If we choose

$$\eta_0 = \frac{\int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p-1} \left((p-1)u_{\varepsilon} + q \ln \frac{1}{\varepsilon} \right) \psi_1 dx}{\int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p-1} \left((p-1)u_{\varepsilon} + q \ln \frac{1}{\varepsilon} \right) \psi_2 dx} > 0,$$

then by the chain rule, $t'(0) = -\frac{D_s(0,1)}{D_t(0,1)} = 0$, which implies that for s small $t(s) = 1 + O(s^2)$. We calculate $I_{\varepsilon}(t(s)\tilde{w}_{\varepsilon})$. Since $supp(\psi_1) \cap supp(\psi_2) = \emptyset$, we obtain

$$\begin{split} &\int_{\Omega} (K_H(x) \nabla \tilde{w}_{\varepsilon} | \nabla \tilde{w}_{\varepsilon}) dx \\ &= \int_{\Omega} (K_H(x) \nabla u_{\varepsilon} | \nabla u_{\varepsilon}) dx + s^2 \int_{\Omega} (K_H(x) \nabla \psi_1 | \nabla \psi_1) dx + \eta_0^2 s^2 \int_{\Omega} (K_H(x) \nabla \psi_2 | \nabla \psi_2) dx \\ &+ 2s \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_+^p \psi_1 dx - 2s \eta_0 \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_+^p \psi_2 dx. \end{split}$$

Since $t(s) = 1 + O(s^2)$, we have

$$\frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(t(s)\tilde{w}_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx$$

$$= \frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(t(s)u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx + \frac{1}{\varepsilon^{2}} \int_{\Omega} (t(s)u_{\varepsilon} - q \ln \frac{1}{\varepsilon})_{+}^{p} (s\psi_{1} - \eta_{0}s\psi_{2}) dx$$

$$+ \frac{p}{2\varepsilon^{2}} \int_{\Omega} \left(t(s)u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p-1} (s\psi_{1} - \eta_{0}s\psi_{2})^{2} dx + O(s^{2+\sigma})$$

$$= \frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(t(s)u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p+1} dx + \frac{1}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} (s\psi_{1} - \eta_{0}s\psi_{2}) dx$$

$$+ \frac{p}{2\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p-1} (s^{2}\psi_{1}^{2} + \eta_{0}^{2}s^{2}\psi_{2}^{2}) dx + O(s^{2+\sigma}),$$

for some $\sigma > 0$. So by Proposition 2.2, we get

$$c_{\varepsilon} \leq I_{\varepsilon}(t(s)\tilde{w}_{\varepsilon})$$

$$= \frac{t(s)^{2}}{2} \int_{\Omega} (K_{H}(x)\nabla\tilde{w}_{\varepsilon}|\nabla\tilde{w}_{\varepsilon})dx - \frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(t(s)\tilde{w}_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_{+}^{p+1} dx$$

$$= \frac{t(s)^{2}}{2} \int_{\Omega} (K_{H}(x)\nabla u_{\varepsilon}|\nabla u_{\varepsilon})dx - \frac{1}{(p+1)\varepsilon^{2}} \int_{\Omega} \left(t(s)u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_{+}^{p+1} dx$$

$$+ \frac{s^{2}}{2} \left[\int_{\Omega} (K_{H}(x)\nabla\psi_{1}|\nabla\psi_{1})dx - \frac{p}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_{+}^{p-1} \psi_{1}^{2}dx + \eta_{0}^{2} \left(\int_{\Omega} (K_{H}(x)\nabla\psi_{2}|\nabla\psi_{2})dx - \frac{p}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_{+}^{p-1} \psi_{2}^{2}dx \right) \right] + O(s^{2+\sigma}).$$

Note that

$$\frac{t(s)^2}{2} \int_{\Omega} (K_H(x) \nabla u_{\varepsilon} | \nabla u_{\varepsilon}) dx - \frac{1}{(p+1)\varepsilon^2} \int_{\Omega} \left(t(s) u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_+^{p+1} dx \le \max_{t \ge 0} I_{\varepsilon}(t u_{\varepsilon}).$$

Direct calculation shows that

$$\int_{\Omega} (K_H(x)\nabla\psi_1|\nabla\psi_1)dx - \frac{p}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_+^{p-1} \psi_1^2 dx$$

$$= \int_{A_1} \left(K_H(x)\nabla\left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)|\nabla\psi_1\right) dx - \frac{p}{\varepsilon^2} \int_{A_1} \left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)_+^{p-1} \psi_1^2 dx$$

$$= -\frac{p-1}{\varepsilon^2} \int_{A_1} (\psi_1)^{p+1} dx < 0.$$

Here we have used $\mathcal{L}q = 0$. Similarly, $\int_{\Omega} (K_H(x) \nabla \psi_2 | \nabla \psi_2) dx - \frac{p}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_+^{p-1} \psi_2^2 dx < 0$. Thus $c_{\varepsilon} < \max_{t \geq 0} I_{\varepsilon}(tu_{\varepsilon}) = I_{\varepsilon}(u_{\varepsilon}) = c_{\varepsilon}$, which is clearly a contradiction. So we conclude that A_{ε} is connected.

Moreover, we can prove that A_{ε} is simply connected. Let U be the connected component of $\Omega \setminus A_{\varepsilon}$ such that $\partial \Omega \subseteq \overline{U}$. Note that $\Omega \setminus U$ is open and $\mathcal{L}\left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon}\right) \geq 0$ in $\Omega \setminus U$ and $u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \geq 0$ on $\partial(\Omega \setminus U)$. By the strong maximum principle, $u_{\varepsilon} - q \ln \frac{1}{\varepsilon} > 0$ in $\Omega \setminus U$. Thus A_{λ} is simply connected.

Finally, by the definition of capacity and $u_{\varepsilon} \geq q \ln \frac{1}{\varepsilon}$ on A_{ε} , we have

$$\left(\ln\frac{1}{\varepsilon}\right)^{2} \int_{\Omega \setminus A_{\varepsilon}} q^{2} \left(K_{H}(x) \nabla \left(\frac{u_{\varepsilon}}{q \ln\frac{1}{\varepsilon}}\right) | \nabla \left(\frac{u_{\varepsilon}}{q \ln\frac{1}{\varepsilon}}\right) \right) dx \ge \left(\ln\frac{1}{\varepsilon}\right)^{2} \inf_{\Omega} (q^{2} \lambda_{2}) \int_{\Omega \setminus A_{\varepsilon}} \left| \nabla \left(\frac{u_{\varepsilon}}{q \ln\frac{1}{\varepsilon}}\right) \right|^{2} dx \\
\ge C \left(\ln\frac{1}{\varepsilon}\right)^{2} \lambda_{2} \operatorname{cap}(A_{\varepsilon}, \Omega),$$

where $\lambda_2(x) = \frac{k^2}{k^2 + |x|^2}$ is the smaller eigenvalue of K_H . Since $\mathbb{R}^2 \setminus \Omega$ is connected and unbounded, by the classical estimates of capacity (see [11, 24]), we have

$$\operatorname{cap}(A_{\varepsilon}, \Omega) \ge \frac{2\pi}{\ln 16 \left(1 + \frac{2\operatorname{dist}(A_{\varepsilon}, \partial\Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)}.$$

By Lemmas 2.3, 2.4 and Proposition 3.1,

$$\left(\ln\frac{1}{\varepsilon}\right)^2 \int_{\Omega \setminus A_{\varepsilon}} q^2 \left(K_H(x) \nabla \left(\frac{u_{\varepsilon}}{q \ln\frac{1}{\varepsilon}}\right) |\nabla \left(\frac{u_{\varepsilon}}{q \ln\frac{1}{\varepsilon}}\right)\right) dx \le \int_{\Omega} (K_H(x) \nabla u_{\varepsilon} |\nabla u_{\varepsilon}) dx$$

$$\le \frac{2(p+1)}{p-1} I_{\varepsilon}(u_{\varepsilon}) \le c \ln\frac{1}{\varepsilon}.$$

Combining all these inequalities, we get $\lim_{\varepsilon \to 0} \frac{diam(A_{\varepsilon})}{dist(A_{\varepsilon},\partial\Omega)} = 0$.

Define the energy of the vortex core $E_c(\varepsilon) = \int_{A_{\varepsilon}} \left(K_H(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx$. We will show that $E_c(\varepsilon)$ is uniformly bounded with respect to ε .

Lemma 3.3. There holds for some C independent of ε

$$\int_{A_{\varepsilon}} \left(K_H(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx \le C.$$

Proof. Direct calculation yields that

$$\int_{A_{\varepsilon}} \left(K_H(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx = \frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^{p+1} dx, \quad (3.12)$$

and

$$\int_{\Omega} (K_H(x)\nabla u_{\varepsilon}|\nabla u_{\varepsilon})dx - \int_{A_{\varepsilon}} \left(K_H(x)\nabla\left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)|\nabla\left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)\right)dx \\
= \frac{\ln\frac{1}{\varepsilon}}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q\ln\frac{1}{\varepsilon}\right)^p qdx. \tag{3.13}$$

By (3.12), (K2) and the classical Gagliardo-Nirenberg inequality, we get

$$\begin{split} &\int_{A_{\varepsilon}} \left(K_{H}(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx \\ \leq & \frac{C^{*}}{\varepsilon^{2}} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^{p} dx \left(\int_{A_{\varepsilon}} \left| \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right|^{2} dx \right)^{\frac{1}{2}} \\ \leq & \frac{C^{*}}{\sqrt{\inf_{\Omega} \lambda_{2} \varepsilon^{2}}} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^{p} dx \left(\int_{A_{\varepsilon}} \left(K_{H}(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx \right)^{\frac{1}{2}}, \\ \text{which implies that} \end{split}$$

$$\int_{A_{\varepsilon}} \left(K_H(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx \le C \left(\frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^p dx \right)^2.$$

By (3.13) and Proposition 3.1, we get

$$\frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^p dx \le \frac{C}{\ln \frac{1}{\varepsilon}} \int_{\Omega} (K_H(x) \nabla u_{\varepsilon} | \nabla u_{\varepsilon}) dx \le \frac{C}{\ln \frac{1}{\varepsilon}} I_{\varepsilon}(u_{\varepsilon}) \le C.$$

Thus we get

$$\int_{A_{\varepsilon}} \left(K_H(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx = \frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^{p+1} dx \le C.$$

Using Lemma 3.3, we can get the lower bound of the diameter of the vortex core A_{ε} as follows.

Lemma 3.4. There exists a constant $R_1 > 0$ independent of ε such that

$$diam(A_{\varepsilon}) \ge R_1 \varepsilon.$$

Proof. By (3.12) and the Sobolev inequality, we have

$$\int_{A_{\varepsilon}} \left(K_{H}(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx = \frac{1}{\varepsilon^{2}} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^{p+1} dx$$

$$\leq \frac{C_{*} |A_{\varepsilon}|}{\varepsilon^{2}} \left(\int_{A_{\varepsilon}} \left| \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right|^{2} dx \right)^{\frac{p+1}{2}}$$

$$\leq \frac{C_{*} |A_{\varepsilon}|}{\left(\inf_{\Omega} \lambda_{2}\right)^{\frac{p+1}{2}} \varepsilon^{2}} \left(\int_{A_{\varepsilon}} \left(K_{H}(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx \right)^{\frac{p+1}{2}},$$

from which we deduce that

$$\frac{|A_{\varepsilon}|}{\varepsilon^2} \ge C \left(\int_{A_{\varepsilon}} \left(K_H(x) \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right) \right) dx \right)^{\frac{-p+1}{2}}.$$

By Lemma 3.3, we conclude that $|A_{\varepsilon}| \geq C\varepsilon^2$. Thus we complete the proof by using the isoperimetric inequality $|A_{\varepsilon}| \leq \pi \operatorname{diam}(A_{\varepsilon})^2/4$.

3.3. Asymptotic location of A_{ε} . It follows from Proposition 3.2 that $\lim_{\varepsilon \to 0} A_{\varepsilon} = 0$, that is, the vortex core of u_{ε} will shrink to a single point \hat{x} as $\varepsilon \to 0$. We now prove that the limiting location of A_{ε} is a minimum point of $q^2 \sqrt{\det(K_H)}$, by choosing test functions suitably and using the classical stream-function method.

Proposition 3.5. There holds

$$\lim_{\varepsilon \to 0} dist(A_{\varepsilon}, \hat{x}) = 0,$$

where \hat{x} is a minimizer of $q^2\sqrt{\det(K_H)}$. As a consequence, there holds

$$\lim_{\varepsilon \to 0} \frac{c_{\varepsilon}}{\ln \frac{1}{\varepsilon}} = \pi \min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)}.$$
 (3.14)

Proof. It follows from (3.13) that

$$\frac{\ln\frac{1}{\varepsilon}}{\varepsilon^{2}} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon} \right)^{p} q dx$$

$$= \int_{\Omega} (K_{H}(x)\nabla u_{\varepsilon}|\nabla u_{\varepsilon}) dx - \int_{A_{\varepsilon}} \left(K_{H}(x)\nabla\left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon}\right) |\nabla\left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon}\right)\right) dx$$

$$= 2I_{\varepsilon}(u_{\varepsilon}) - \frac{p-1}{(p+1)\varepsilon^{2}} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon} \right)^{p+1} dx.$$

Hence by Lemma 3.3, we get

$$\frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^p q dx = \frac{2c_{\varepsilon}}{\ln \frac{1}{\varepsilon}} + O\left(\frac{1}{\ln \frac{1}{\varepsilon}} \right) \le C. \tag{3.15}$$

For any $0 < \tau < \sigma \le 1$, define $w_{\varepsilon}^{\sigma,\tau} := \min \left\{ \frac{\left(u_{\varepsilon} - q \ln \frac{1}{\sigma}\right)_{+}}{q \ln \frac{\sigma}{\tau}}, 1 \right\} \in H_{0}^{1}(\Omega)$ and $A_{\varepsilon}^{\sigma} := \frac{\left\{x \in \Omega \mid u_{\varepsilon}(x) > q(x) \ln \frac{1}{\sigma}\right\}$. Then one computes directly that $w_{\varepsilon}^{\sigma,\tau} \equiv 1$ on A_{ε}^{τ} and $supp(w_{\varepsilon}^{\sigma,\tau}) = \frac{A_{\varepsilon}^{\sigma}}{A_{\varepsilon}^{\sigma}}$.

We claim that for every $\varepsilon \leq \tau$,

$$\ln \frac{\sigma}{\tau} \int_{\Omega} q^2 (K_H(x) \nabla w_{\varepsilon}^{\sigma,\tau} | \nabla w_{\varepsilon}^{\sigma,\tau}) dx = \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} q dx.$$
 (3.16)

Indeed, multiplying both sides of (2.1) by $\phi = w_{\varepsilon}^{\sigma,\tau} q \in H_0^1(\Omega)$ and using integration by parts, we get

$$\int_{\Omega} (K_H(x) \nabla u_{\varepsilon} | \nabla (w_{\varepsilon}^{\sigma,\tau} q)) dx = \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} w_{\varepsilon}^{\sigma,\tau} q dx.$$

Direct computations show that

$$\frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} w_{\varepsilon}^{\sigma,\tau} q dx = \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} q dx$$

and

$$\int_{\Omega} (K_{H}(x)\nabla u_{\varepsilon}|\nabla(w_{\varepsilon}^{\sigma,\tau}q))dx = \int_{\Omega} \left(K_{H}(x)\nabla\left(u_{\varepsilon} - q\ln\frac{1}{\sigma}\right)|\nabla(w_{\varepsilon}^{\sigma,\tau}q)\right)dx
= \ln\frac{\sigma}{\tau} \int_{\Omega} (K_{H}(x)\nabla(w_{\varepsilon}^{\sigma,\tau}q)|\nabla(w_{\varepsilon}^{\sigma,\tau}q))dx + \int_{\Omega} \left(K_{H}(x)\nabla\left(u_{\varepsilon} - q\ln\frac{1}{\sigma} - \ln\frac{\sigma}{\tau}w_{\varepsilon}^{\sigma,\tau}q\right)|\nabla(w_{\varepsilon}^{\sigma,\tau}q)\right)dx
= \ln\frac{\sigma}{\tau} \int_{\Omega} (K_{H}(x)\nabla(w_{\varepsilon}^{\sigma,\tau}q)|\nabla(w_{\varepsilon}^{\sigma,\tau}q))dx + \int_{A_{\varepsilon}^{\tau}} \left(K_{H}(x)\nabla\left(u_{\varepsilon} - q\ln\frac{1}{\tau}\right)|\nabla q\right)dx
= \ln\frac{\sigma}{\tau} \int_{\Omega} (K_{H}(x)\nabla(w_{\varepsilon}^{\sigma,\tau}q)|\nabla(w_{\varepsilon}^{\sigma,\tau}q))dx
= \ln\frac{\sigma}{\tau} \int_{\Omega} q^{2}(K_{H}(x)\nabla w_{\varepsilon}^{\sigma,\tau}|\nabla w_{\varepsilon}^{\sigma,\tau}|dx,$$

where we have used the assumption $\mathcal{L}_H q = 0$ and Lemma 2.4. Thus we get (3.16). By the definition of capacity, (3.15) and (3.16), we get

$$\operatorname{cap}(A_{\varepsilon}^{\tau},\Omega) \leq \int_{\Omega} |\nabla w_{\varepsilon}^{1,\tau}|^{2} dx \leq \frac{1}{\inf_{\Omega} q^{2} \lambda_{2}} \int_{\Omega} q^{2} (K_{H}(x) \nabla w_{\varepsilon}^{1,\tau} |\nabla w_{\varepsilon}^{1,\tau}) dx \leq \frac{C}{\ln \frac{1}{\tau}}.$$

Using the capacity estimates in [11] again, we get

$$\frac{2\pi}{\ln 16 \left(1 + \frac{2\operatorname{dist}(A_{\varepsilon}^{\tau}, \partial\Omega)}{\operatorname{diam}(A_{\varepsilon}^{\tau})}\right)} \le \operatorname{cap}(A_{\varepsilon}^{\tau}, \Omega),$$

from which we deduce that,

$$\frac{2\pi}{\ln 16 \left(1 + \frac{2\operatorname{dist}(A_{\varepsilon}^{\tau}, \partial\Omega)}{\operatorname{diam}(A_{\varepsilon}^{\tau})}\right)} \le \frac{C}{\ln \frac{1}{\tau}}.$$
(3.17)

So there exist constants $C_1, C_2 > 0$ independent of ε, τ , such that for any $0 < \tau < 1$ and $\varepsilon \leq \tau$,

$$\operatorname{diam}(A_{\varepsilon}^{\tau}) \le C_1 \tau^{C_2}.$$

We now claim that for any $\delta > 0$, there exist $\rho > 0$ and $0 < \varepsilon_0 < \rho$, such that for any $\varepsilon \in (0, \varepsilon_0)$ and $x, y \in A_{\varepsilon}^{\rho}$,

$$q(x)^2 \le q(y)^2 (1+\delta),$$
 (3.18)

and

$$(K_H(x)\zeta|\zeta) \le (1+\delta)(K_H(y)\zeta|\zeta), \quad \forall \ \zeta \in \mathbb{R}^2.$$
(3.19)

Indeed, since $\inf_{\Omega} q > 0$ and $q \in C^2(\Omega) \cap C^1(\overline{\Omega})$, it is easy to get (3.18). By $(\mathcal{K}2)$ and the regularity of K_H , one can also get (3.19).

Thus taking $\sigma = \rho, \tau = \varepsilon$ in (3.16), we get for any $x_{\varepsilon} \in A_{\varepsilon} \subseteq A_{\varepsilon}^{\rho}$

$$\frac{1}{\varepsilon^2 \ln \frac{\rho}{\varepsilon}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} q dx = \int_{\Omega} q^2 (K_H(x) \nabla w_{\varepsilon}^{\rho, \varepsilon} | \nabla w_{\varepsilon}^{\rho, \varepsilon}) dx \ge \frac{q^2 (x_{\varepsilon})}{(1 + \delta)^2} \int_{\Omega} (K_H(x_{\varepsilon}) \nabla w_{\varepsilon}^{\rho, \varepsilon} | \nabla w_{\varepsilon}^{\rho, \varepsilon}) dx.$$
(3.20)

Define a linear transformation matrix $T_{\varepsilon}: \mathbb{R}^2 \to \mathbb{R}^2$ satisfying

$$T_{\varepsilon}K_H(x_{\varepsilon})T_{\varepsilon}^t = Id.$$

Then $|\det(T_{\varepsilon})| = |\det(K_H)(x_{\varepsilon})|^{-\frac{1}{2}}$. Let $\Omega' = T_{\varepsilon}(\Omega)$, $A'_{\varepsilon} = T_{\varepsilon}(A_{\varepsilon})$. For any $y = T_{\varepsilon}(x) \in \Omega'$, define $\bar{w}_{\varepsilon}^{\rho,\varepsilon}(y) = w_{\varepsilon}^{\rho,\varepsilon}(x) = w_{\varepsilon}^{\rho,\varepsilon}(T_{\varepsilon}^{-1}(y))$. Since $\bar{w}_{\varepsilon}^{\rho,\varepsilon} \equiv 1$ on A'_{ε} and $\bar{w}_{\varepsilon}^{\rho,\varepsilon} \in H_0^1(\Omega')$, using the capacity estimates again we have

$$\int_{\Omega} (K_{H}(x_{\varepsilon}) \nabla w_{\varepsilon}^{\rho,\varepsilon} | \nabla w_{\varepsilon}^{\rho,\varepsilon}) dx = \int_{\Omega'} (T_{\varepsilon} K_{H}(x_{\varepsilon}) T_{\varepsilon}^{t} \nabla \bar{w}_{\varepsilon}^{\rho,\varepsilon} | \nabla \bar{w}_{\varepsilon}^{\rho,\varepsilon}) |det(T_{\varepsilon}^{-1})| dy$$

$$= \sqrt{\det(K_{H})} (x_{\varepsilon}) \int_{\Omega'} |\nabla \bar{w}_{\varepsilon}^{\rho,\varepsilon}|^{2} dy$$

$$\geq \sqrt{\det(K_{H})} (x_{\varepsilon}) \cdot \exp(A_{\varepsilon}', \Omega')$$

$$\geq \sqrt{\det(K_{H})} (x_{\varepsilon}) \cdot \frac{2\pi}{\ln 16 \left(1 + \frac{2\operatorname{dist}(A_{\varepsilon}', \partial\Omega')}{\operatorname{diam}(A_{\varepsilon}')}\right)}$$

$$\geq \sqrt{\det(K_{H})} (x_{\varepsilon}) \cdot \frac{2\pi}{\ln 16 \left(1 + \frac{C_{0}\operatorname{dist}(A_{\varepsilon}, \partial\Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)}, \tag{3.21}$$

for some $C_0 > 0$ independent of ε . Thus by (3.20) and (3.21), we get

$$q^{2}\sqrt{\det(K_{H})}(x_{\varepsilon}) \leq \frac{(1+\delta)^{2}}{2\pi \ln \frac{\rho}{\varepsilon}} \ln 16 \left(1 + \frac{C_{0}\operatorname{dist}(A_{\varepsilon}, \partial\Omega)}{\operatorname{diam}(A_{\varepsilon})}\right) \left(\frac{1}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon}\right)_{+}^{p} q dx\right). \tag{3.22}$$

Taking the limit superior in both sides of (3.22), using (3.15) and Proposition 3.1, we obtain that for any $\delta > 0$ and $x_{\varepsilon} \in A_{\varepsilon}$,

$$\limsup_{\varepsilon \to 0} q^2 \sqrt{\det(K_H)}(x_{\varepsilon}) \le \frac{(1+\delta)^2}{2\pi} \limsup_{\varepsilon \to 0} \frac{\ln 16 \left(1 + \frac{C_0 \operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)}{\ln \frac{\rho}{\varepsilon}} \cdot 2\pi \min_{x \in \overline{\Omega}} q^2 \sqrt{\det(K_H)}(x).$$

By Lemma 3.4 and $\operatorname{dist}(A_{\varepsilon}, \partial \Omega) \leq \operatorname{diam}(\Omega)$, we get

$$\limsup_{\varepsilon \to 0} \frac{\ln 16 \left(1 + \frac{C_0 \operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})} \right)}{\ln \frac{\rho}{\varepsilon}} \le \limsup_{\varepsilon \to 0} \frac{\ln 16 \left(1 + \frac{C}{\varepsilon} \right)}{\ln \frac{\rho}{\varepsilon}} = 1.$$

Hence we have

$$\limsup_{\varepsilon \to 0} q^2 \sqrt{\det(K_H)}(x_{\varepsilon}) \le (1+\delta)^2 \min_{x \in \overline{\Omega}} q^2 \sqrt{\det(K_H)}(x).$$

By the arbitrariness of $\delta > 0$, we conclude that for any $x_{\varepsilon} \in A_{\varepsilon}$, x_{ε} tends to \hat{x} , where \hat{x} is a minimizer of $q^2 \sqrt{det(K_H)}$.

Taking the limit inferior in both sides of (3.22), using (3.15), and by the the arbitrariness of $\delta > 0$, we have

$$\liminf_{\varepsilon \to 0} \frac{c_{\varepsilon}}{\ln \frac{1}{\varepsilon}} \ge \pi \min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)}.$$

Combining this with Proposition 3.1, we get (3.14). The proof is thus complete.

We can then get estimates of the diameter of A_{ε} as follows.

Lemma 3.6. There holds

$$\lim_{\varepsilon \to 0} \frac{\ln \frac{dist(A_{\varepsilon}, \partial \Omega)}{diam(A_{\varepsilon})}}{\ln \frac{1}{\varepsilon}} = 1.$$
 (3.23)

Proof. On the one hand, by Lemma 3.4, we have

$$\limsup_{\varepsilon \to 0} \frac{\ln \frac{\operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})}}{\ln \frac{1}{\varepsilon}} \leq \limsup_{\varepsilon \to 0} \frac{\ln \frac{C}{\varepsilon}}{\ln \frac{1}{\varepsilon}} = 1.$$

On the other hand, taking the limit inferior in both sides of (3.22) and using (3.15) and Proposition 3.1, we get

$$\frac{(1+\delta)^2}{2\pi} \liminf_{\varepsilon \to 0} \frac{\ln 16 \left(1 + \frac{C_0 \operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)}{\ln \frac{\rho}{\varepsilon}} \cdot 2\pi \min_{x \in \overline{\Omega}} q^2 \sqrt{\det(K_H)}(x)$$

$$\geq \liminf_{\varepsilon \to 0} q^2 \sqrt{\det(K_H)}(x_{\varepsilon})$$

$$= \min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)},$$

which implies that

$$\frac{1}{(1+\delta)^2} \leq \liminf_{\varepsilon \to 0} \frac{\ln 16 \left(1 + \frac{C_0 \operatorname{dist}(A_\varepsilon, \partial \Omega)}{\operatorname{diam}(A_\varepsilon)}\right)}{\ln \frac{\rho}{\varepsilon}} = \liminf_{\varepsilon \to 0} \frac{\ln \frac{\operatorname{dist}(A_\varepsilon, \partial \Omega)}{\operatorname{diam}(A_\varepsilon)}}{\ln \frac{1}{\varepsilon}}.$$

By the arbitrariness of $\delta > 0$, we have $\liminf_{\varepsilon \to 0} \frac{\ln \frac{\operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})}}{\ln \frac{1}{\varepsilon}} \geq 1$. The proof is thus complete.

Remark 3.7. A direct consequence of Lemmas 3.6 and 3.4 is that for any $\alpha \in (0,1)$, there exists $C_1, C_2 > 0$ such that

$$C_1 \varepsilon \leq \operatorname{diam}(A_{\varepsilon}) \leq C_2 \varepsilon^{\alpha}.$$

When the limiting location \bar{x} of A_{ε} is on the boundary of Ω , such an estimate is optimal. Similar results have been found for 2D Euler equations and 3D axisymmetric equations, see [11, 20, 24] for example. However when $\bar{x} \in \Omega$, we can improve estimates of A_{ε} .

By Proposition 3.5, we show that the limiting location of A_{ε} is x^* , where $q^2 \sqrt{\det(K_H)}(x^*) = \min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)}$. Note that $\kappa(w_{\varepsilon}) = \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon}\right)_+^p dx$ is the circulation of $w_{\varepsilon} = \frac{1}{\varepsilon^2} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon}\right)_+^p$. The limit of $\kappa(w_{\varepsilon})$ can be obtained as follows.

Lemma 3.8. There holds

$$\lim_{\varepsilon \to 0} \kappa(w_{\varepsilon}) = 2\pi q \sqrt{\det(K_H)}(x^*).$$

Proof. It follows from (3.13), (3.12) and the definition of c_{ε} that

$$\frac{\ln\frac{1}{\varepsilon}}{\varepsilon^{2}} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon} \right)^{p} q dx = \int_{\Omega} (K_{H} \nabla u_{\varepsilon} | \nabla u_{\varepsilon}) dx - \int_{A_{\varepsilon}} \left(K_{H} \nabla \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon} \right) | \nabla \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon} \right) \right) dx$$

$$= 2c_{\varepsilon} - \frac{p-1}{(p+1)\varepsilon^{2}} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon} \right)^{p+1} dx. \tag{3.24}$$

By Lemma 3.3, we have

$$\frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^{p+1} dx \le C.$$

So $\frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^p q dx = \frac{2c_{\varepsilon}}{\ln \frac{1}{\varepsilon}} + O\left(\frac{1}{\ln \frac{1}{\varepsilon}} \right)$. By Proposition 3.5 and the fact that $A_{\varepsilon} \to x^*$, we get

$$\lim_{\varepsilon \to 0} \kappa(w_{\varepsilon}) = q(x^*)^{-1} \lim_{\varepsilon \to 0} \frac{1}{\varepsilon^2} \int_{A_{\varepsilon}} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)^p q dx = q(x^*)^{-1} \lim_{\varepsilon \to 0} \frac{2c_{\varepsilon}}{\ln \frac{1}{\varepsilon}} = 2\pi q(x^*) \sqrt{\det(K_H(x^*))}.$$

3.4. Further analysis when the limiting location of A_{ε} is in Ω . When the limiting location of A_{ε} is in Ω , we can improve the results in Proposition 3.5 and Lemma 3.6 by giving more accurate estimates of lower bound of c_{ε} and upper bound of the diameter of A_{ε} . Indeed, we have

Proposition 3.9. If for any $x_{\varepsilon} \in A_{\varepsilon}$, $\lim_{\varepsilon \to 0} x_{\varepsilon} = x^* \in \Omega$, then

$$c_{\varepsilon} = I_{\varepsilon}(u_{\varepsilon}) = \pi \min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)} \ln \frac{1}{\varepsilon} + O(1).$$
 (3.25)

Moreover, there exist $R_1, R_2 > 0$ such that

$$R_1 \varepsilon \le diam(A_{\varepsilon}) \le R_2 \varepsilon.$$

Proof. By (3.17), we get for any $0 < \tau < 1$ and $\varepsilon \le \tau$,

$$\frac{2\pi}{\ln 16 \left(1 + \frac{2\operatorname{dist}(A_{\varepsilon}^{\tau}, \partial\Omega)}{\operatorname{diam}(A_{\varepsilon}^{\tau})}\right)} \leq \frac{C}{\ln \frac{1}{\tau}}.$$

So there exist $C_0, \alpha_0 > 0$ such that $\frac{\operatorname{dist}(A_{\varepsilon}^{\tau}, \partial\Omega)}{\operatorname{diam}(A_{\varepsilon}^{\tau})} \geq \frac{C_0}{\tau^{\alpha_0}}$, which implies that $\operatorname{diam}(A_{\varepsilon}^{\tau}) \leq C_1 \tau^{\alpha_0}$ for some $C_1 > 0$. That is, for any $x, y \in A_{\varepsilon}^{\tau}$, $|x - y| \leq C_1 \tau^{\alpha_0}$.

By $q \in C^2(\Omega) \cap C^1(\overline{\Omega})$, $(K_H)_{ij} \in C^{\infty}(\Omega)$ for i, j = 1, 2, $\inf_{\Omega} q > 0$ and $(\mathcal{K}2)$, similarly to the proof of (3.18) and (3.19), we can get that q^2K_H is Dini-continuous uniformly in Ω ,

which means that, there exists a non-negative function $\gamma(s)$, such that $\int_0^{s_0} \frac{\gamma(s)}{s} ds < +\infty$ for some $s_0 > 0$ and

$$q^{2}(x)(K_{H}(x)\zeta|\zeta) \leq (1 + \gamma(|x - y|))q^{2}(y)(K_{H}(y)\zeta|\zeta), \quad \forall x, y \in \Omega, \quad \zeta \in \mathbb{R}^{2}.$$
 (3.26)

Thus by (3.16) and (3.26), we get for any $0 < \tau < \sigma < 1$, $\varepsilon \le \tau$ and $x_{\varepsilon} \in A_{\varepsilon} \subseteq A_{\varepsilon}^{\sigma}$,

$$\frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} q dx = \ln \frac{\sigma}{\tau} \int_{\Omega} q^2 (K_H(x) \nabla w_{\varepsilon}^{\sigma, \tau} | \nabla w_{\varepsilon}^{\sigma, \tau}) dx$$

$$\geq \ln \frac{\sigma}{\tau} \frac{1}{1 + \gamma(C_1 \sigma^{\alpha_0})} \int_{\Omega} q^2 (x_{\varepsilon}) (K_H(x_{\varepsilon}) \nabla w_{\varepsilon}^{\sigma, \tau} | \nabla w_{\varepsilon}^{\sigma, \tau}) dx,$$

which implies that

$$\left(\ln\frac{\sigma}{\tau}\right)^2 \int_{\Omega} q^2(x_{\varepsilon}) (K_H(x_{\varepsilon}) \nabla w_{\varepsilon}^{\sigma,\tau} | \nabla w_{\varepsilon}^{\sigma,\tau}) dx \le \ln\frac{\sigma}{\tau} (1 + \gamma (C_1 \sigma^{\alpha_0})) \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon}\right)_{+}^{p} q dx.$$
(3.27)

Taking $\varepsilon = \tau_1 < \sigma_1 = \tau_2 < \sigma_2 = \tau_3 < \dots < \sigma_k = \rho$, and summing (3.27) over $\{j = 1, 2, \dots, k\}$, we get for any $x_{\varepsilon} \in A_{\varepsilon}$

$$\left(\ln\frac{\rho}{\varepsilon}\right)^2 \int_{\Omega} q^2(x_{\varepsilon}) (K_H(x_{\varepsilon}) \nabla w_{\varepsilon}^{\rho,\varepsilon} | \nabla w_{\varepsilon}^{\rho,\varepsilon}) dx \leq \left(\ln\frac{\rho}{\varepsilon} + \sum_{j=1}^k \gamma(C_1 \sigma_j^{\alpha_0}) \ln\frac{\sigma_j}{\tau_j}\right) \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon}\right)_+^p q dx.$$

By taking the limit of Riemann sums in the above inequality, we have

$$\left(\ln\frac{\rho}{\varepsilon}\right)^2 \int_{\Omega} q^2(x_{\varepsilon}) (K_H(x_{\varepsilon}) \nabla w_{\varepsilon}^{\rho,\varepsilon} | \nabla w_{\varepsilon}^{\rho,\varepsilon}) dx \le \left(\ln\frac{\rho}{\varepsilon} + \int_{\varepsilon}^{\rho} \frac{\gamma(C_1 \sigma^{\alpha_0})}{\sigma} d\sigma\right) \frac{1}{\varepsilon^2} \int_{\Omega} \left(u_{\varepsilon} - q \ln\frac{1}{\varepsilon}\right)_{+}^{\rho} q dx.$$

Since

$$\int_{\varepsilon}^{\rho} \frac{\gamma(C_1 \sigma^{\alpha_0})}{\sigma} d\sigma = \int_{\varepsilon^{\alpha_0}}^{\rho^{\alpha_0}} \frac{\gamma(C_1 \sigma')}{\sigma'^{\frac{1}{\alpha_0}}} \frac{1}{\alpha_0} \sigma'^{\frac{1}{\alpha_0} - 1} d\sigma' = \frac{1}{\alpha_0} \int_{C_1 \varepsilon^{\alpha_0}}^{C_1 \rho^{\alpha_0}} \frac{\gamma(\sigma')}{\sigma'} d\sigma' < +\infty,$$

we get

$$\int_{\Omega} q^{2}(x_{\varepsilon}) (K_{H}(x_{\varepsilon}) \nabla w_{\varepsilon}^{\rho,\varepsilon} | \nabla w_{\varepsilon}^{\rho,\varepsilon}) dx \leq \left(\frac{1}{\ln \frac{\rho}{\varepsilon}} + \frac{C}{\left(\ln \frac{\rho}{\varepsilon}\right)^{2}} \right) \frac{1}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon} \right)_{+}^{p} q dx, \quad (3.28)$$

which is the refined version of (3.20). So repeating the proof of Proposition 3.5, we have

$$q^{2}\sqrt{\det(K_{H})}(x_{\varepsilon}) \leq \frac{\ln 16\left(1 + \frac{C_{0}\operatorname{dist}(A_{\varepsilon},\partial\Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)}{2\pi} \left(\frac{1}{\ln \frac{\rho}{\varepsilon}} + \frac{C}{\left(\ln \frac{\rho}{\varepsilon}\right)^{2}}\right) \left(\frac{1}{\varepsilon^{2}} \int_{\Omega} \left(u_{\varepsilon} - q \ln \frac{1}{\varepsilon}\right)_{+}^{p} q dx\right),\tag{3.29}$$

which improves (3.22).

Thus, taking (3.15) into (3.29) and using Proposition 3.1, we obtain

$$\min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)} \ln \frac{1}{\varepsilon} \leq \frac{\ln 16 \left(1 + \frac{C_0 \operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)}{2\pi} \left(\frac{1}{\ln \frac{\rho}{\varepsilon}} + \frac{C}{\left(\ln \frac{\rho}{\varepsilon}\right)^2}\right) \left(2c_{\varepsilon} + O(1)\right) \\
\leq \frac{\ln 16 \left(1 + \frac{C_0 \operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)}{2\pi} \left(\frac{1}{\ln \frac{\rho}{\varepsilon}} + \frac{C}{\left(\ln \frac{\rho}{\varepsilon}\right)^2}\right) \left(2\pi \min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)} \ln \frac{1}{\varepsilon} + O(1)\right).$$

Direct computation shows that

$$\ln \frac{C_0 \operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})} \ge \ln \frac{1}{\varepsilon} + O(1),$$

which implies that $diam(A_{\varepsilon}) \leq R_2 \varepsilon$ for some $R_2 > 0$.

Finally, by taking diam $(A_{\varepsilon}) \geq R_1 \varepsilon$ (see Lemma 3.4) into (3.29) and using (3.15), we get

$$c_{\varepsilon} \geq \pi \min_{\overline{\Omega}} q^{2} \sqrt{\det(K_{H})} \frac{1}{\ln 16 \left(1 + \frac{C_{0} \operatorname{dist}(A_{\varepsilon}, \partial \Omega)}{\operatorname{diam}(A_{\varepsilon})}\right)} \frac{1}{\left(\frac{1}{\ln \frac{\rho}{\varepsilon}} + \frac{C}{\left(\ln \frac{\rho}{\varepsilon}\right)^{2}}\right)} \ln \frac{1}{\varepsilon} + O(1)$$

$$\geq \pi \min_{\overline{\Omega}} q^{2} \sqrt{\det(K_{H})} \frac{1}{\ln \frac{C}{\varepsilon}} \frac{1}{\left(\frac{1}{\ln \frac{\rho}{\varepsilon}} + \frac{C}{\left(\ln \frac{\rho}{\varepsilon}\right)^{2}}\right)} \ln \frac{1}{\varepsilon} + O(1)$$

$$\geq \pi \min_{\overline{\Omega}} q^{2} \sqrt{\det(K_{H})} \ln \frac{1}{\varepsilon} + O(1).$$

Combining this with Proposition 3.1, we get (3.25). The proof is thus complete.

4. Proof of Theorem 1.5 and 1.2

4.1. **Proof of Theorem 1.5.** In subsections 3.1 to 3.4, we prove the existence and asymptotic behavior of solutions of (2.1) under the additional assumption that there exist minimum points of $q^2 \sqrt{\det(K_H)}$ on the x_1 -axis.

Now we give proof of Theorem 1.5 in the case that all minimum points of $q^2\sqrt{\det(K_H)}$ is not on the x_1 -axis. Let $\hat{x} = (|\hat{x}|\cos\theta_{\hat{x}}, |\hat{x}|\sin\theta_{\hat{x}})$ be a minimizer of $q^2\sqrt{\det(K_H)}$ on $\overline{\Omega}$.

Let $\Omega_{\hat{x}} = \{\bar{R}_{\hat{x}}x \mid x \in \Omega\}$. For any function $u \in \mathcal{H}(\Omega)$, let $u_{\hat{x}}(x) = u(\bar{R}_{-\hat{x}}x)$ for any $x \in \Omega_{\hat{x}}$. So $u_{\hat{x}} \in \mathcal{H}(\Omega_{\hat{x}})$. Let $q_{\hat{x}}(x) = q(\bar{R}_{-\hat{x}}x)$ for any $x \in \Omega_{\hat{x}}$. Then by Lemma 2.6, we get that u is a solution of (2.1) if and only if $u_{\hat{x}}$ is a solution of

$$\begin{cases} -\operatorname{div}(K_H(x)\nabla v) = \frac{1}{\varepsilon^2} \left(v - q_{\hat{x}} \ln \frac{1}{\varepsilon} \right)_+^p, & \text{in } \Omega_{\hat{x}}. \\ v = 0, & \text{on } \partial \Omega_{\hat{x}}. \end{cases}$$
(4.1)

For equations (4.1), we claim that there exist minimum points of $q_{\hat{x}}^2 \sqrt{\det(K_H)}$ on the x_1 -axis. Indeed, we can prove that $(|\hat{x}|, 0) = \bar{R}_{\hat{x}} \hat{x} \in \overline{\Omega_{\hat{x}}}$ and

$$q_{\hat{x}}^2 \sqrt{\det(K_H)}(\bar{R}_{\hat{x}}\hat{x}) = \frac{kq_{\hat{x}}^2(\bar{R}_{\hat{x}}\hat{x})}{\sqrt{k^2 + |\bar{R}_{\hat{x}}\hat{x}|^2}} = \frac{kq^2(\hat{x})}{\sqrt{k^2 + |\hat{x}|^2}} = \min_{\overline{\Omega}} q^2 \sqrt{\det(K_H)} = \min_{\overline{\Omega}_{\hat{x}}} q_{\hat{x}}^2 \sqrt{\det(K_H)},$$

which implies that $(|\hat{x}|, 0)$ is a minimizer of $q_{\hat{x}}^2 \sqrt{\det(K_H)}$.

Hence we can repeat the proof in subsections 2.1 and 3.1-3.4 to show that there exist a family of solutions v_{ε} of (4.1) concentrating near minimum points of $q_{\hat{x}}^2 \sqrt{\det(K_H)}$. Define $u_{\varepsilon}(x) = v_{\varepsilon}(\bar{R}_{\hat{x}}x)$ for any $x \in \Omega$. Then $\{u_{\varepsilon}\}$ is a family of solutions of (2.1) which concentrates near a minimizer of $q^2 \sqrt{\det(K_H)}$ as $\varepsilon \to 0$.

Let $w_{\varepsilon} = \mathcal{L}_H u_{\varepsilon}$, $\varphi_{\varepsilon} = u_{\varepsilon} - q \ln \frac{1}{\varepsilon}$. Then $(w_{\varepsilon}, \varphi_{\varepsilon})$ is the desired solution pair of Theorem 1.5.

4.2. **Proof of Theorem 1.2.** Based on results in Theorem 1.5, we now give the proof of Theorem 1.2. Let q(x) = m for every m > 0. Then q satisfies $\mathcal{L}_H q = 0$. So by Theorem 1.5, there exist solutions u_{ε} of (1.18) with $f(t) = t_+^p$ and $\mu = m \ln \frac{1}{\varepsilon}$ concentrating near x^* , which is a minimizer of $q^2 \sqrt{\det(K_H)}$. Since

$$q^2 \sqrt{\det(K_H)}(x) = m^2 \left(\frac{k^2}{k^2 + |x|^2}\right)^{\frac{1}{2}},$$

we get that $x^* \in \overline{\Omega}$ satisfies $|x^*| = \max_{\overline{\Omega}} |x|$. Let $w_{\varepsilon} = \mathcal{L}_H u_{\varepsilon}$. By Lemma 3.8, the limit of circulation is

$$\lim_{\varepsilon \to 0} \kappa(w_{\varepsilon}) = 2\pi q(x^*) \sqrt{\det(K_H(x^*))} = 2\pi m \cdot \left(\frac{k^2}{k^2 + |x^*|^2}\right)^{\frac{1}{2}} = \frac{2k\pi m}{\sqrt{k^2 + |x^*|^2}}.$$

To conclude, $(w_{\varepsilon}, u_{\varepsilon})$ is the desired solution pair and the proof of Theorem 1.2 is complete.

Remark 4.1. From the proof of Theorem 1.2, we see that the limiting location x^* of w_{ε} satisfies $|x^*| = \max_{\overline{\Omega}} |x|$. So x^* must be on the boundary of Ω . This implies that results in subsection 3.4 can not hold. In this case, the optimal estimates of diameter of A_{ε} is Lemma 3.6, rather than Proposition 3.9.

Acknowledgments: D. Cao was supported by NNSF of China (grant No. 11831009) and Chinese Academy of Sciences by grant QYZDJ-SSW-SYS021. J. Wan was supported by NNSF of China (grant No. 12101045) and Beijing Institute of Technology Research Fund Program for Young Scholars (No.3170011182016).

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